

# Dark sector as origin of tiny lepton masses and new sources of $(g-2)_\mu$

Kei Yagyu (Osaka U.)

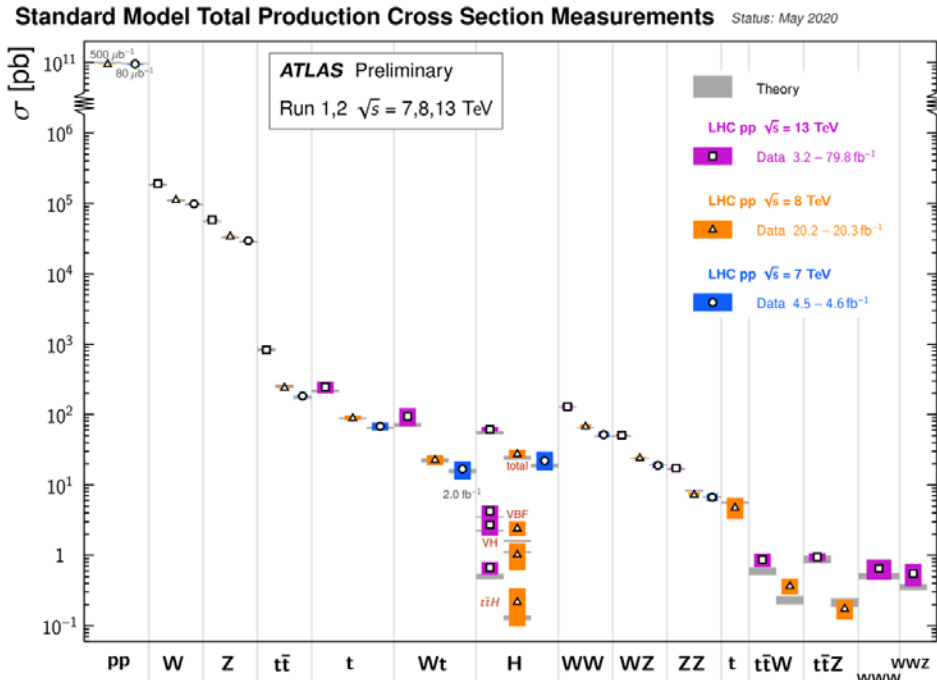


*Kai-Feng Chen, Cheng-Wei Chiang, KY, 2006.07929 [hep-ph] (JHEP)*

*Cheng-Wei Chiang, KY, 2104.00890 [hep-ph] (PRD)*

12<sup>th</sup> October, Nagoya U. (Online)

# SM well describes high energy phenomena



SM predictions are good agreement.

## ATLAS Exotics Searches\* - 95% CL Upper Exclusion Limits

Status: May 2020

ATLAS Preliminary  
Status: May 2020  
 $\int \mathcal{L} dt = (3.2 - 139) \text{ fb}^{-1}$   
 $\sqrt{s} = 8, 13 \text{ TeV}$

	Model	$\ell, \gamma$	Jets <sup>†</sup>	$E_{\text{T}}^{\text{miss}}$	$\int \mathcal{L} dt [\text{fb}^{-1}]$	Limit	Reference
Extra dimensions	ADD $G_{KK} + g/q$	$0, e, \mu$	$1 - 4 j$	Yes	36.1	$M_{\text{Pl}}$	1711.03301
	ADD non-resonant $\gamma\gamma$	$2 \gamma$	-	-	36.7	$M_{\text{Pl}}$	1707.04147
	ADD QBH	-	$2 j$	-	37.0	$M_{\text{Pl}}$	1703.09127
	ADD BH high $\Sigma p_T$	$\geq 1, e, \mu$	$\geq 2 j$	-	3.2	$M_{\text{Pl}}$	1606.02265
	ADD BH multiplet	-	$\geq 3 j$	-	3.6	$M_{\text{Pl}}$	1512.02586
	RS1 $G_{KK} \rightarrow \gamma\gamma$	$2 \gamma$	-	-	36.7	$G_{KK} \text{ mass}$	1707.04147
	Bulk RS $G_{KK} \rightarrow WW/ZZ$	multi-channel	-	-	36.1	$G_{KK} \text{ mass}$	1608.02380
	Bulk RS $G_{KK} \rightarrow WV \rightarrow \ell\nu qq$	$1, e, \mu$	$2 j / 1 j$	Yes	139	$G_{KK} \text{ mass}$	2004.14636
	Bulk RS $G_{KK} \rightarrow tt$	$1, e, \mu$	$\geq 1 b, \geq 1 j$	Yes	36.1	$G_{KK} \text{ mass}$	1804.10823
	2UED / RPP	$1, e, \mu$	$\geq 2 b, \geq 3 j$	Yes	36.1	$KK \text{ mass}$	1803.09678
Gauge bosons	SSM $Z' \rightarrow \ell\ell$	$2, e, \mu$	-	-	139	$Z' \text{ mass}$	1903.06248
	SSM $Z' \rightarrow \tau\tau$	$2, \tau$	-	-	36.1	$Z' \text{ mass}$	1709.07242
	Leptophobic $Z' \rightarrow bb$	-	$2 b$	-	36.1	$Z' \text{ mass}$	1805.08299
	Leptophobic $Z' \rightarrow \tau\tau$	$0, e, \mu$	$\geq 1 b, \geq 2 j$	Yes	139	$Z' \text{ mass}$	2005.05138
	SSM $W' \rightarrow \ell\nu$	$1, e, \mu$	-	Yes	139	$W' \text{ mass}$	1906.05609
	SSM $W' \rightarrow \tau\nu$	$1, \tau$	-	Yes	36.1	$W' \text{ mass}$	1801.06992
	HVT $W' \rightarrow WZ \rightarrow \ell\nu qq$ model B	$1, e, \mu$	$2 j / 1 j$	Yes	139	$W' \text{ mass}$	2004.14636
	HVT $V' \rightarrow WV \rightarrow qq qq$ model B	$0, e, \mu$	$2 j$	-	139	$V' \text{ mass}$	1906.08589
	HVT $V' \rightarrow WH/ZH$ model B	multi-channel	-	-	139	$V' \text{ mass}$	1712.06518
	HVT $W' \rightarrow WH$ model B	multi-channel	-	-	139	$V' \text{ mass}$	1906.08589
LRSB $W_R \rightarrow \mu N_R$	multi-channel	-	-	36.1	$W_R \text{ mass}$	1807.10473	
	$2 \mu$	$1 j$	-	80	$W_R \text{ mass}$	1904.12679	
CI	CI $qqqq$	-	$2 j$	-	37.0	$A$	1703.09127
	CI $\ell\ell qq$	$2, e, \mu$	-	-	139	$A$	1703.09127
	CI $tttt$	$\geq 1, e, \mu$	$\geq 1 b, \geq 1 j$	Yes	36.1	$A$	1811.02305
						$2.57 \text{ TeV}$	$\eta_{\text{CI}}$
DM	Axial-vector mediator (Dirac DM)	$0, e, \mu$	$1 - 4 j$	Yes	36.1	$m_{\text{mediator}}$	1711.03301
	Colored scalar mediator (Dirac DM)	$0, e, \mu$	$1 - 4 j$	Yes	36.1	$m_{\text{mediator}}$	1711.03301
	VV <sub>XY</sub> EFT (Dirac DM)	$0, e, \mu$	$1, 1, \leq 1 j$	Yes	3.2	$M_{\text{Pl}}$	1606.02272
	Scalar reson. $\phi \rightarrow \tau\tau$ (Dirac DM)	$0, 1, e, \mu$	$1 b, 0, 1 j$	Yes	36.1	$m_{\text{Pl}}$	1812.08743
LQ	Scalar LQ 1 <sup>st</sup> gen	$1, 2, e, \mu$	$\geq 2 j$	Yes	36.1	$LQ \text{ mass}$	1902.00377
	Scalar LQ 2 <sup>nd</sup> gen	$1, 2, \mu$	$\geq 2 j$	Yes	36.1	$LQ \text{ mass}$	1902.00377
	Scalar LQ 3 <sup>rd</sup> gen	$2, \tau$	$2 b$	-	36.1	$LQ \text{ mass}$	1902.08103
	Scalar LQ 3 <sup>rd</sup> gen	$0, 1, e, \mu$	$2 b$	Yes	36.1	$LQ \text{ mass}$	1902.08103
Heavy quarks	VLO $TT \rightarrow Ht/Zt/Wb + X$	multi-channel	-	-	36.1	$T \text{ mass}$	1808.02343
	VLO $BB \rightarrow Wt/Zb + X$	multi-channel	-	-	36.1	$B \text{ mass}$	1807.11883
	VLO $T_{3/2} T_{3/2} \rightarrow Wt + X$	$2(SS) \geq 3, e, \mu \geq 1 b, \geq 1 j$	Yes	36.1	$T_{3/2} \text{ mass}$	1812.07343	
	VLO $Y \rightarrow Wb + X$	$1, e, \mu$	$\geq 1 b, \geq 1 j$	Yes	36.1	$Y \text{ mass}$	1812.07343
	VLO $B \rightarrow Hb + X$	$0, e, \mu, 2 \gamma$	$\geq 1 b, \geq 1 j$	Yes	79.8	$B \text{ mass}$	1509.04261
	VLO $QQ \rightarrow WqWq$	$1, e, \mu$	$\geq 4 j$	Yes	20.3	$Q \text{ mass}$	1509.04261
Excited fermions	Excited quark $q^* \rightarrow qg$	-	$2 j$	-	139	$q^* \text{ mass}$	1910.08447
	Excited quark $q^* \rightarrow q\gamma$	$1 \gamma$	$1 j$	-	36.7	$q^* \text{ mass}$	1709.10440
	Excited quark $b^* \rightarrow bg$	-	$1 b, 1 j$	-	36.1	$b^* \text{ mass}$	1805.09299
	Excited lepton $\ell^*$	$3, e, \mu$	-	-	20.3	$\ell^* \text{ mass}$	1411.2921
	Excited lepton $\nu^*$	$3, e, \mu, \tau$	-	-	20.3	$\nu^* \text{ mass}$	1411.2921
Other	Type III Seesaw	$1, e, \mu$	$\geq 2 j$	Yes	79.8	$N \text{ mass}$	1809.11105
	LRSB Majorana $\nu$	$2, \mu$	$2 j$	-	36.1	$N \text{ mass}$	1710.09748
	Higgs triplet $H^{\pm\pm} \rightarrow \ell\ell$	$2, 3, 4, e, \mu (SS)$	-	-	36.1	$H^{\pm\pm} \text{ mass}$	1411.2921
	Higgs triplet $H^{\pm\pm} \rightarrow \ell\tau$	$3, e, \mu, \tau$	-	-	20.3	$H^{\pm\pm} \text{ mass}$	1411.2921
	Multi-charged particles	-	-	-	36.1	multi-charged particle mass	1812.07343
	Magnetic monopoles	-	-	-	34.4	monopole mass	1905.10130
		$\sqrt{s} = 8 \text{ TeV}$	$\sqrt{s} = 13 \text{ TeV}$ partial data	$\sqrt{s} = 13 \text{ TeV}$ full data			
				$10^{-1}$	1	10	Mass scale [TeV]

\*Only a selection of the available mass limits on new states or phenomena is shown.

†Small-radius (large-radius) jets are denoted by the letter j (J).

New particles have not been observed.

Q. Does the SM enough?

A. Of course, **No!**

# New physics must exist

## Established BSM Phenomena

- Neutrino oscillations
- Dark matter, dark energy
- Baryon asymmetry of the Universe
- Cosmic inflation


## Theoretical Issues

- Origin of EWSB
- Flavor hierarchies
- Strong QCD
- Unification of the forces
- Quantization of gravity, ...

Some of these problems suggest the existence of **TeV scale NP**,  
but why are they not found so far??

# Where is New Physics?

1. New particles are heavy.

2. New particles are light, but ... 

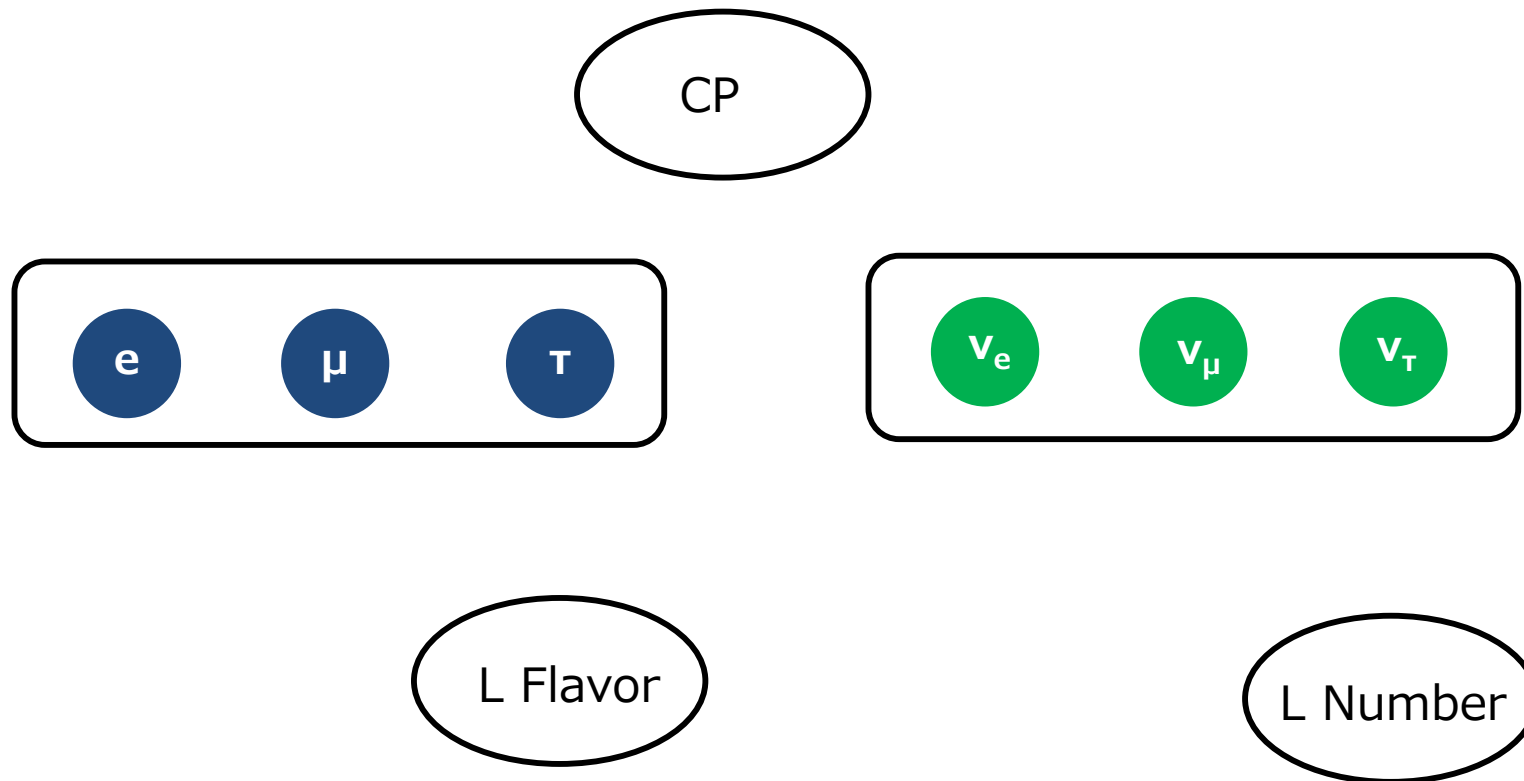
- they feebly couple to SM particles.
- their decay is hidden by backgrounds.

Q. Are there any “**traces**” of NP effects in low energy observables?

A. Yes.

**Lepton sector** is a good example to provide such observables.

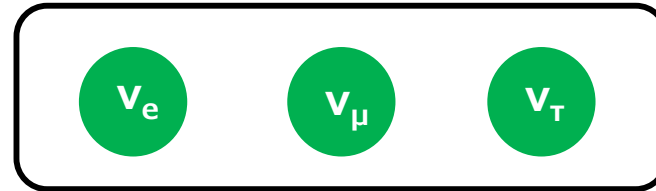
# Lepton Sector Physics



# Lepton Sector Physics

$$\mathcal{L}_{\text{eff}} = \mathcal{L}_{\text{SM}} + \frac{c}{\Lambda} L_L L_L H H + \frac{c'}{\Lambda^2} (\bar{L}_L \sigma^{\mu\nu} i \gamma_5 e_R) F_{\mu\nu} H + \dots$$

~~CP~~

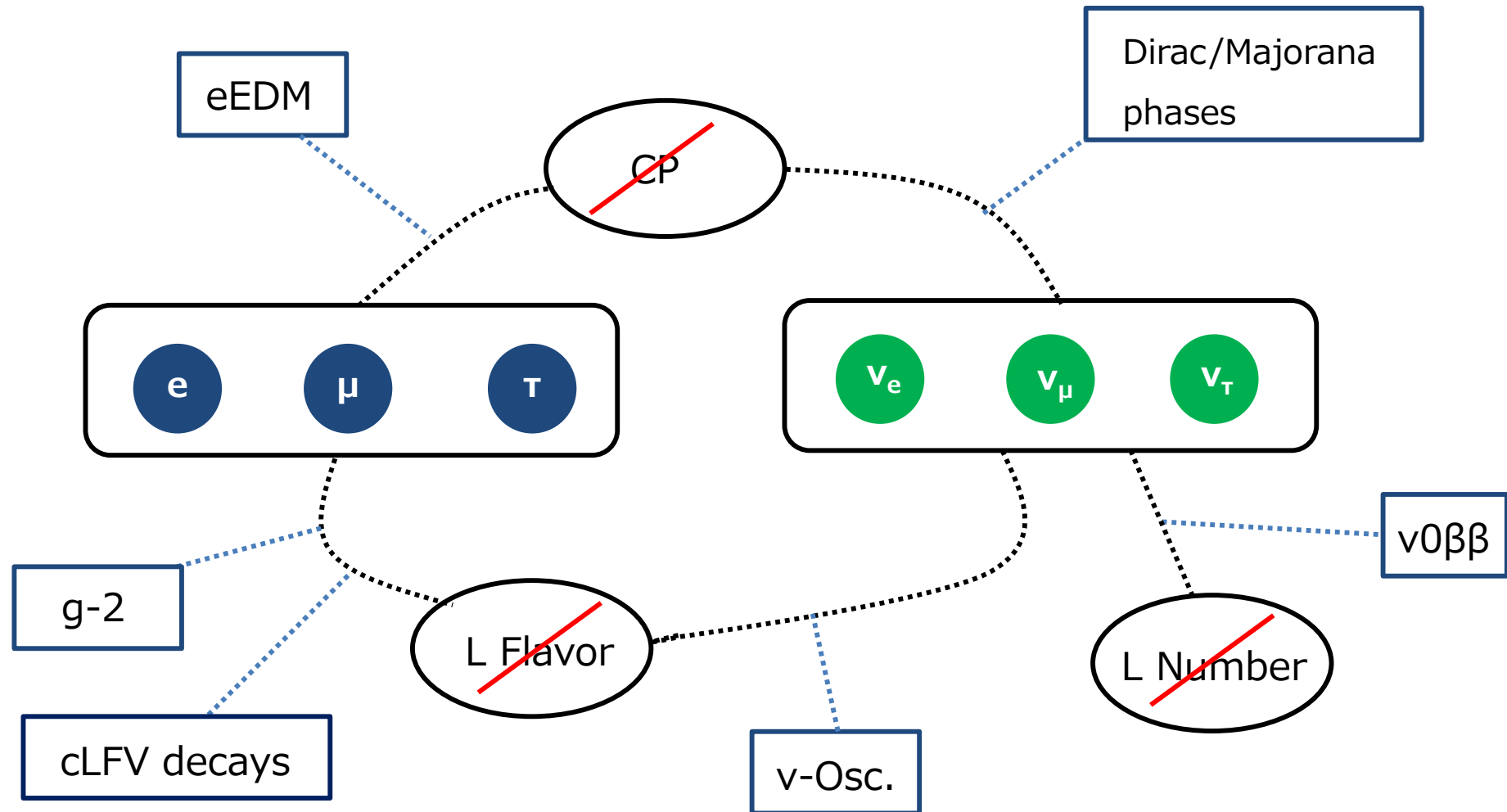


~~L Flavor~~

~~L Number~~

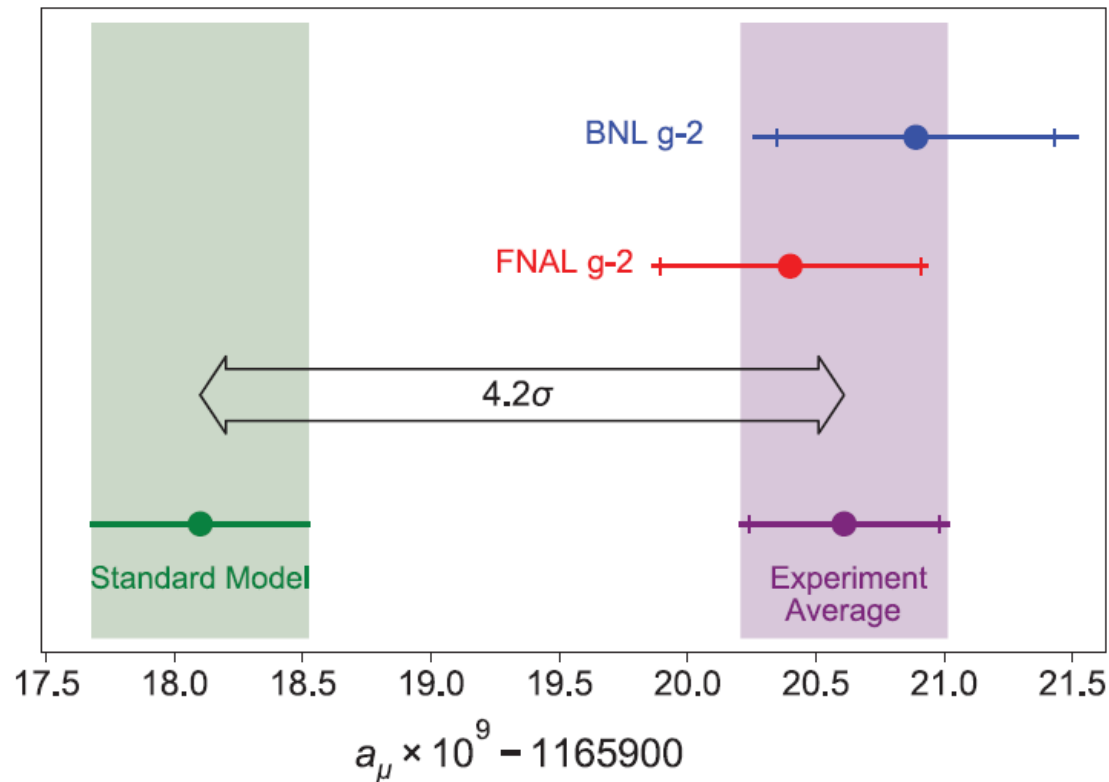
# Lepton Sector Physics

$$\mathcal{L}_{\text{eff}} = \mathcal{L}_{\text{SM}} + \frac{c}{\Lambda} L_L L_L H H + \frac{c'}{\Lambda^2} (\bar{L}_L \sigma^{\mu\nu} i \gamma_5 e_R) F_{\mu\nu} H + \dots$$



# Muon g-2 anomaly

FNAL, 2104.03281 [hep-ex]



$$\Delta a_\mu \equiv a_\mu^{\text{Exp}} - a_\mu^{\text{SM}}$$

Before:  $\Delta a_\mu = (2.79 \pm 0.76) \times 10^{-9} \quad (3.7\sigma)$

After:  $\Delta a_\mu = (2.51 \pm 0.59) \times 10^{-9} \quad (4.2\sigma!!)$



# Festival on arXiv

[16] [arXiv:2104.03217](#) [pdf, other]

**Supersymmetric Interpretation of the Muon  $g - 2$  Anomaly**  
Motoi Endo, Koichi Hamaguchi, Sho Iwamoto, Teppei Kitahara

[17] [arXiv:2104.03223](#) [pdf, other]

**Wino-Higgsino dark matter in the MSSM from the  $g - 2$  anomaly**  
Sho Iwamoto, Tsutomu T. Yanagida, Norimi Yokozaki

[18] [arXiv:2104.03227](#) [pdf, other]

**Lepton-specific inert two-Higgs-doublet model confronted with the new results for muon and electron  $g-2$  anomaly and multi-lepton searches at the LHC**  
Xiao-Fang Han, Tianjun Li, Hong-Xin Wang, Lei Wang, Yang Zhang

[19] [arXiv:2104.03228](#) [pdf, other]

**Muon  $g - 2$  and  $B$ -anomalies from Dark Matter**  
Giorgio Arcadi, Lorenzo Calibbi, Marco Fedele, Federico Mescia

[20] [arXiv:2104.03231](#) [pdf, other]

**Confronting spin-3/2 and other new fermions with the muon  $g-2$  measurement**  
Juan C. Criado, Abdelhak Djouadi, Niko Koivunen, Kristjan Mürsepp, Martti Raidal, Hardi Veermäe

[21] [arXiv:2104.03238](#) [pdf, other]

**Probing light dark matter with scalar mediator: muon  $(g - 2)$  deviation, the proton radius puzzle**  
Bin Zhu, Xuewen Liu

[22] [arXiv:2104.03239](#) [pdf, other]

**Heavy Bino and Slepton for Muon  $g-2$  Anomaly**  
Yuchao Gu, Ning Liu, Liangliang Su, Daohan Wang

[23] [arXiv:2104.03242](#) [pdf, other]

**Revisiting the  $\mu$ - $\tau$ -philic Higgs doublet in light of the muon  $g - 2$  anomaly,  $\tau$  decays, and multi-lepton searches at the**  
Hong-Xin Wang, Lei Wang, Yang Zhang

• • •

[46] [arXiv:2104.03302](#) [pdf, other]

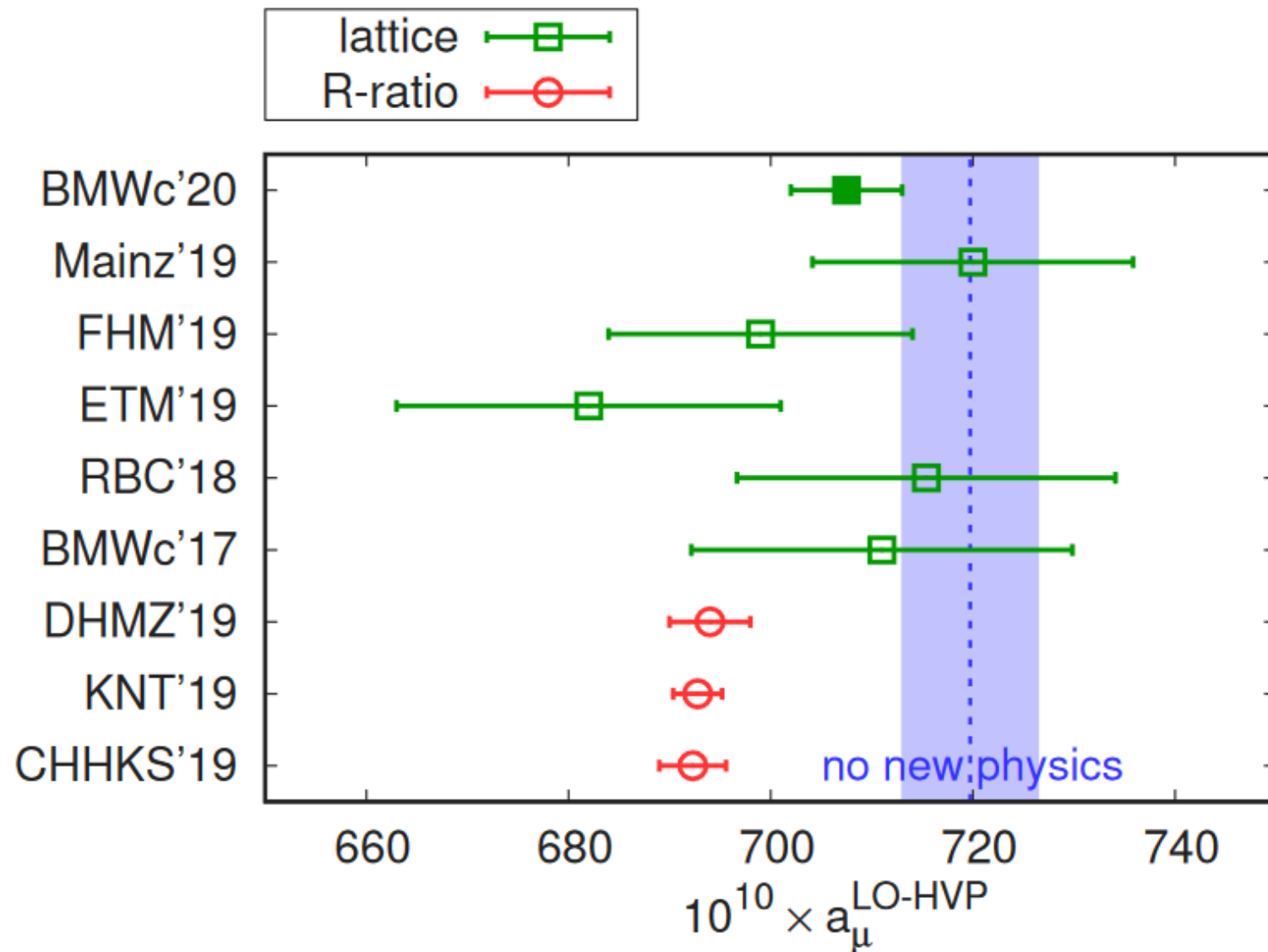
**The Tiny  $(g-2)$  Muon Wobble from Small- $\mu$  Supersymmetry**  
Sebastian Baum, Marcela Carena, Nausheen R. Shah, Carlos E. M. Wagner

31 papers (8<sup>th</sup> April)

- SUSY
- Extended Higgs
- WIMP DM
- Axion-like particles
- Extended gauge sym.
- ...

# Muon $g-2$ anomaly

*BMW Collaboration, 2002.12347 [hep-lat]*



# Electron g-2 anomaly

## Measurement of the fine-structure constant as a test of the Standard Model

Richard H. Parker,<sup>1\*</sup> Chenghui Yu,<sup>1\*</sup> Weicheng Zhong,<sup>1</sup> Brian Estey,<sup>1</sup> Holger Müller<sup>1,2†</sup>

Measurements of the fine-structure constant  $\alpha$  require methods from across subfields and are thus powerful tests of the consistency of theory and experiment in physics. Using the recoil frequency of cesium-133 atoms in a matter-wave interferometer, we recorded the most accurate measurement of the fine-structure constant to date:  $\alpha = 1/137.035999046(27)$  at  $2.0 \times 10^{-10}$  accuracy. Using multiphoton interactions (Bragg diffraction and Bloch oscillations), we demonstrate the largest phase (12 million radians) of any Ramsey-Bordé interferometer and control systematic effects at a level of 0.12 part per billion. Comparison with Penning trap measurements of the electron gyromagnetic anomaly  $g_e - 2$  via the Standard Model of particle physics is now limited by the uncertainty in  $g_e - 2$ ; a 2.5 $\sigma$  tension rejects dark photons as the reason for the unexplained part of the muon's magnetic moment at a 99% confidence level. Implications for dark-sector candidates and electron substructure may be a sign of physics beyond the Standard Model that warrants further investigation.

$$\Delta a_e \equiv a_e^{\text{exp}} - a_e^{\text{SM}} = (-8.8 \pm 3.6) \times 10^{-13} \quad (-2.5 \sigma)$$



$$\Delta a_e = (4.8 \pm 3.0) \times 10^{-13} \quad (1.6\sigma) \quad \text{Nature, 588 (2020)}$$

Science

Science 13 Apr 2018:  
Vol. 360, Issue 6385, pp. 191-195  
DOI: 10.1126/science.aap7706

$$\frac{c}{\Lambda} \bar{\ell}_L \sigma^{\mu\nu} \ell_R F_{\mu\nu}, \quad c \sim \frac{m_\ell}{v}$$

$$\left| \frac{\Delta a_e}{\Delta a_\mu} \right| \sim \frac{1}{3000} \gg \frac{m_e^2}{m_\mu^2} \left( \sim \frac{1}{40000} \right)$$

# CPV in neutrino oscillations

2108.08219 [hep-ex]

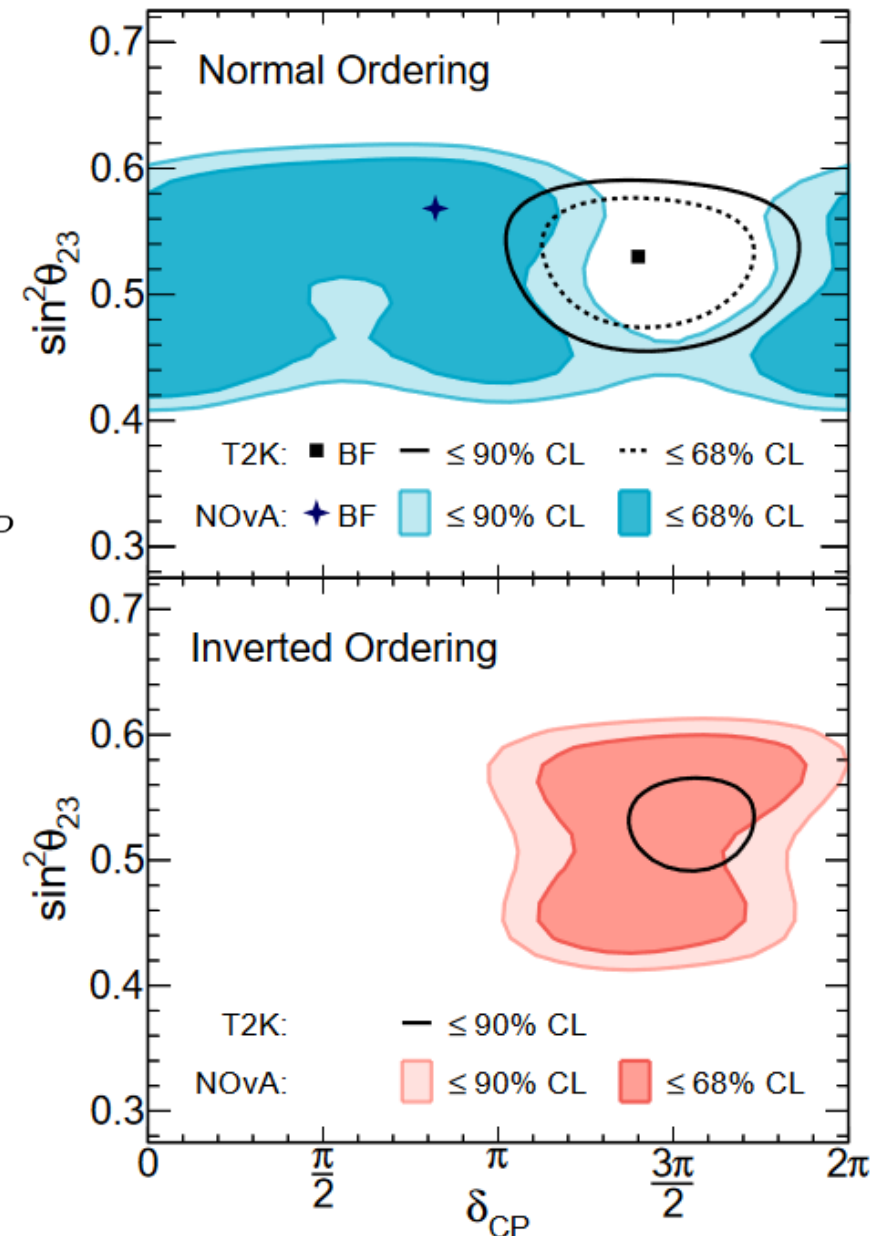
CPV in the neutrino sector can be sizable.

*C. Jarlskog (1985)*

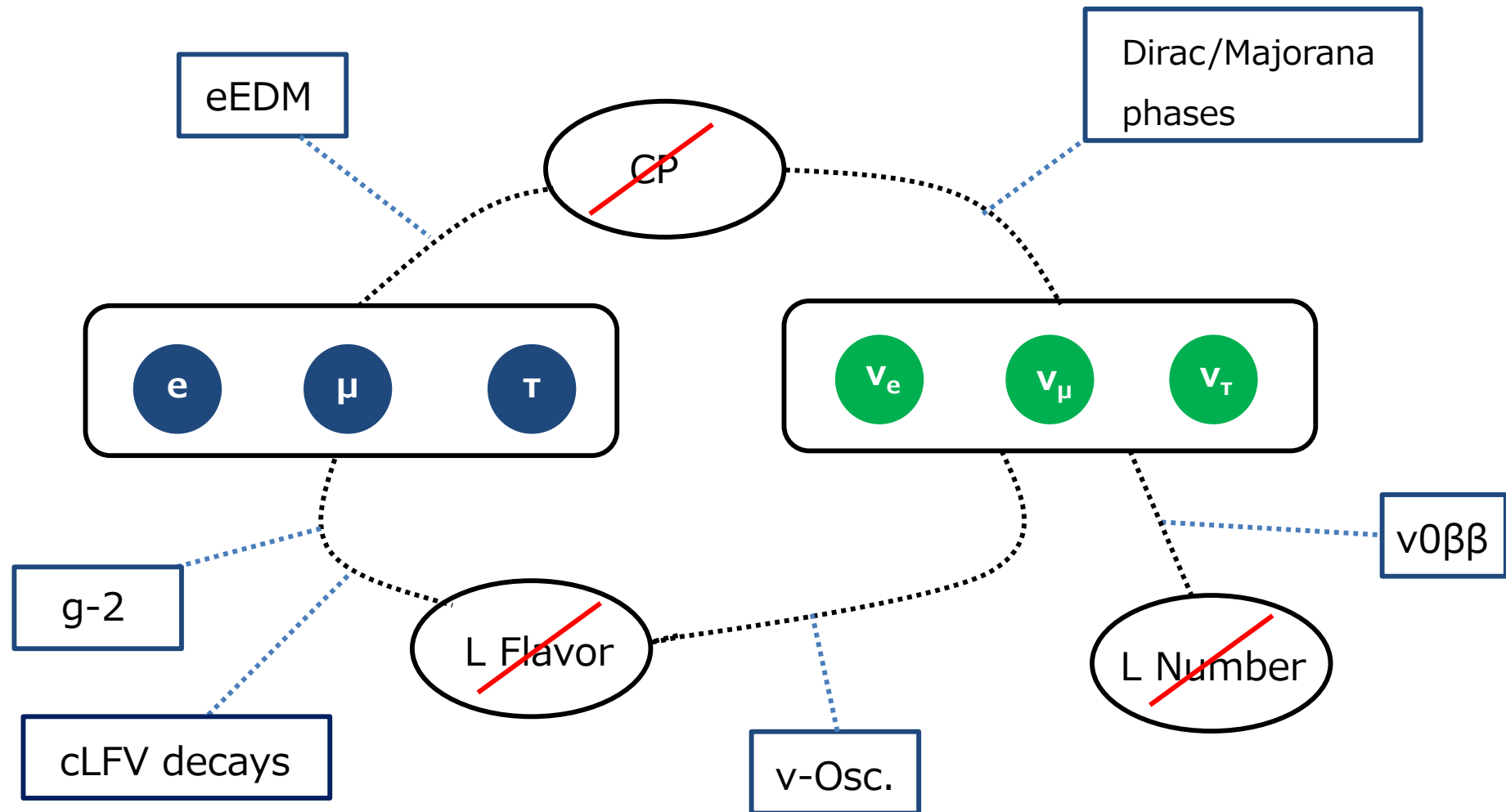
$$J_{CP}^{\text{CKM}} \equiv \text{Im}(V_{us}V_{ub}^*V_{cs}^*V_{cb}) \sim 3 \times 10^{-5}$$

$$J_{CP}^{\text{PMNS}} \equiv \text{Im}(U_{e2}U_{e3}^*U_{\mu 2}^*U_{\mu 3}) \sim 3.3 \times 10^{-2} \times \sin \delta_{CP}$$

Such a sizable CPV can be related to the **baryon asymmetry of the Universe**.



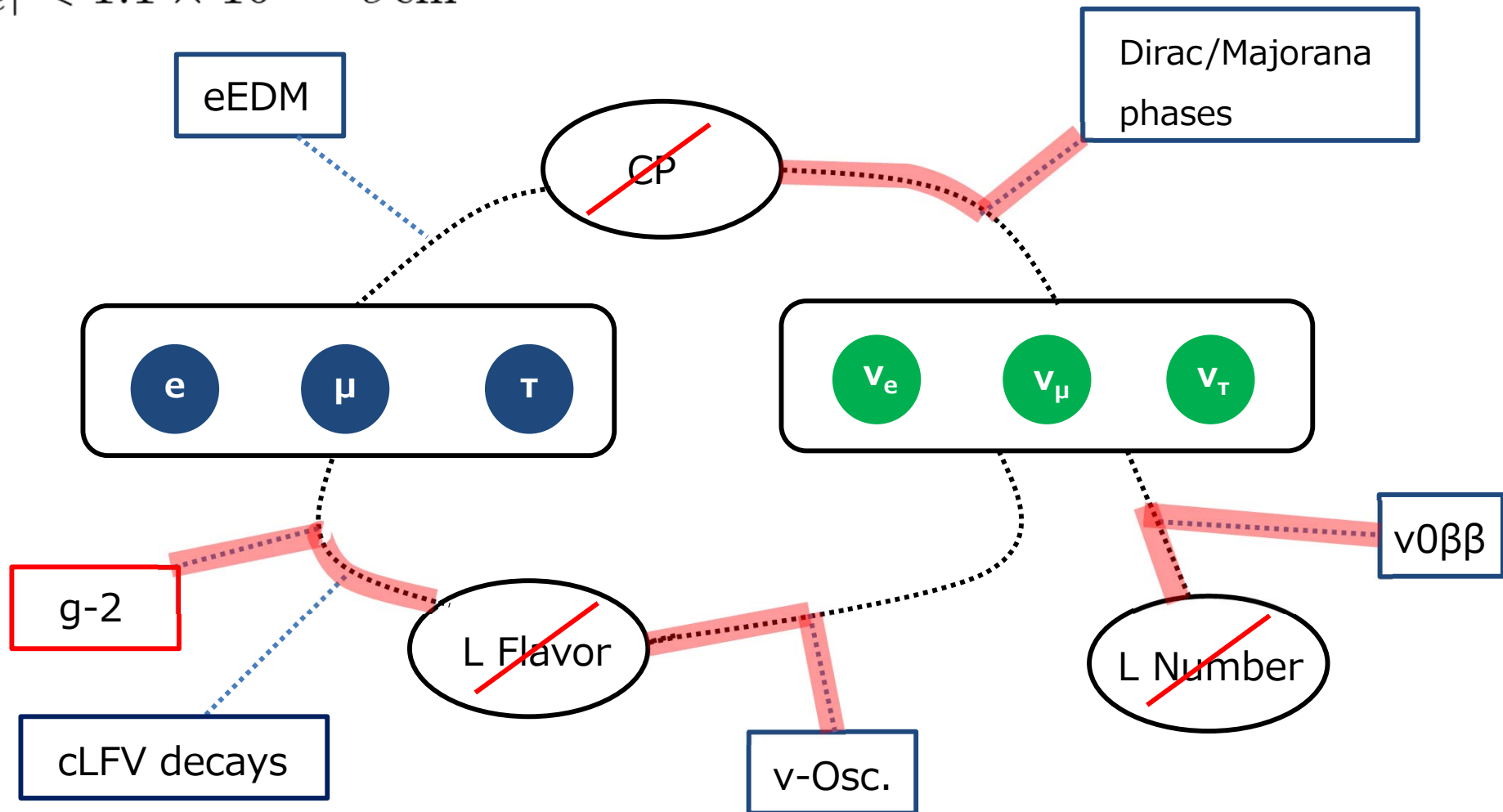
# Lepton Sector Physics



# Lepton Sector Physics

$$|d_e| < 1.1 \times 10^{-29} \text{ e cm}$$

ACME Collaboration (2018)



$$\mathcal{B}(\mu \rightarrow e\gamma) < 4.2 \times 10^{-13}$$

MEG Collaboration (2016)

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I. Introduction

II. Muon  $g-2$  in BSMs

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V. Summary

# New physics contribution at 1-loop

- Effective dim. 5 dipole operator

$$\mathcal{L}_{\text{eff}} = -\frac{1}{2} \left( \frac{e}{2m} a_f \right) \bar{\mu}_L \sigma^{\mu\nu} \mu_R F_{\mu\nu} + \text{h.c.}$$



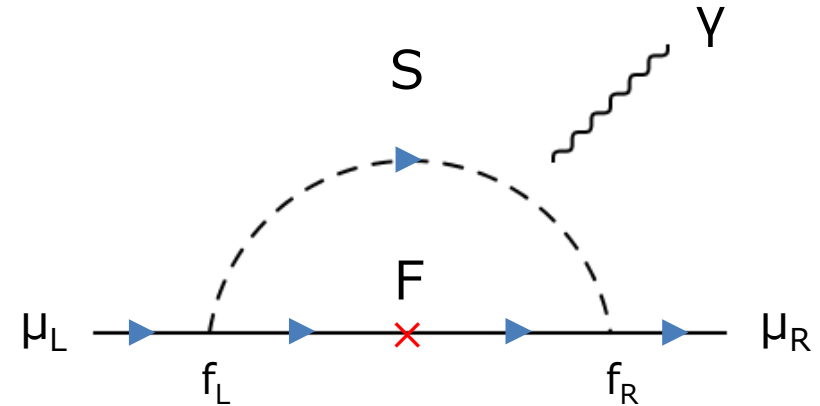
# New physics contribution at 1-loop

- Effective dim. 5 dipole operator

$$\mathcal{L}_{\text{eff}} = -\frac{1}{2} \left( \frac{e}{2m} a_f \right) \bar{\mu}_L \sigma^{\mu\nu} \mu_R F_{\mu\nu} + \text{h.c.}$$

- New interactions:

$$\mathcal{L}_{\text{new}} = \bar{F}(f_L P_L + f_R P_R) \mu S + \text{h.c.}$$



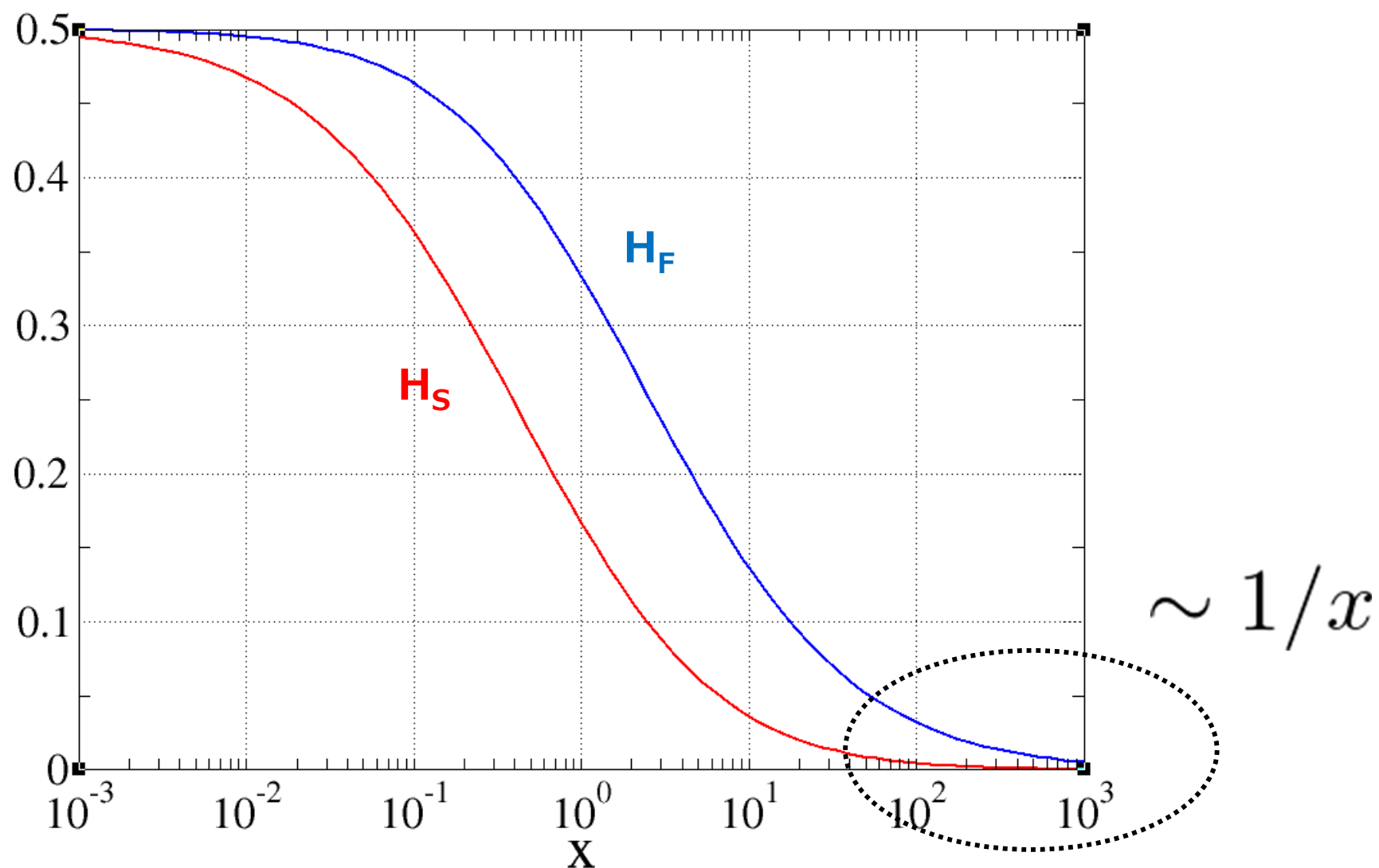
➡

$$a_\mu \simeq \frac{1}{8\pi^2} \frac{m_\mu}{M_F} f_L f_R \left[ -Q_F H_F \left( \frac{M_S^2}{M_F^2} \right) + Q_S H_S \left( \frac{M_S^2}{M_F^2} \right) \right]$$

$$\sim \frac{f_L f_R}{8\pi^2} \left( \frac{m_\mu}{M_F} \right) \left( \frac{M_F}{M_S} \right)^2 \quad \text{for } M_F \ll M_S$$

$$H_F(x) = \frac{1 - 4x + 3x^2 - 2x^2 \ln x}{2(1 - x)^3}$$

$$H_S(x) = \frac{1 - x^2 + 2x \ln x}{2(1 - x)^3}$$



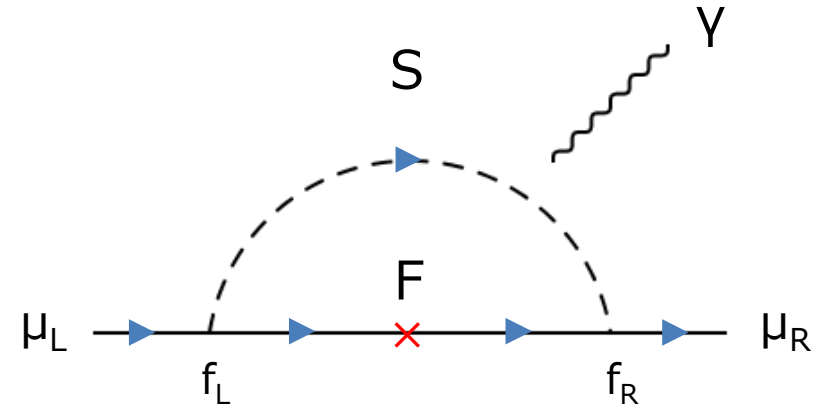
# New physics contribution at 1-loop

- Effective dim. 5 dipole operator

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➡

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$$\sim \frac{f_L f_R}{8\pi^2} \left( \frac{m_\mu}{M_F} \right) \left( \frac{M_F}{M_S} \right)^2 \quad \text{for } M_F \ll M_S$$

There are 2 enhancement sources, i.e., 1) coupling enhancement, 2) chiral enhancement.

Ex. in the charged radiative seesaw model (Sec. 3)

$$a_\mu^{\text{NP}} \sim 3 \times 10^{-9} \times \frac{3 \text{ TeV}}{M} \times \frac{f_L}{0.1} \times \frac{f_R}{0.1}$$

# Quick review of 2 Higgs doublet models

## □ Higgs basis *Davidson, Haber (2005)*

$$\begin{pmatrix} \Phi_1 \\ \Phi_2 \end{pmatrix} = \begin{pmatrix} \cos \beta & -\sin \beta \\ \sin \beta & \cos \beta \end{pmatrix} \begin{pmatrix} \Phi \\ \Phi' \end{pmatrix} \quad \tan \beta = v_2/v_1$$

**NG boson**

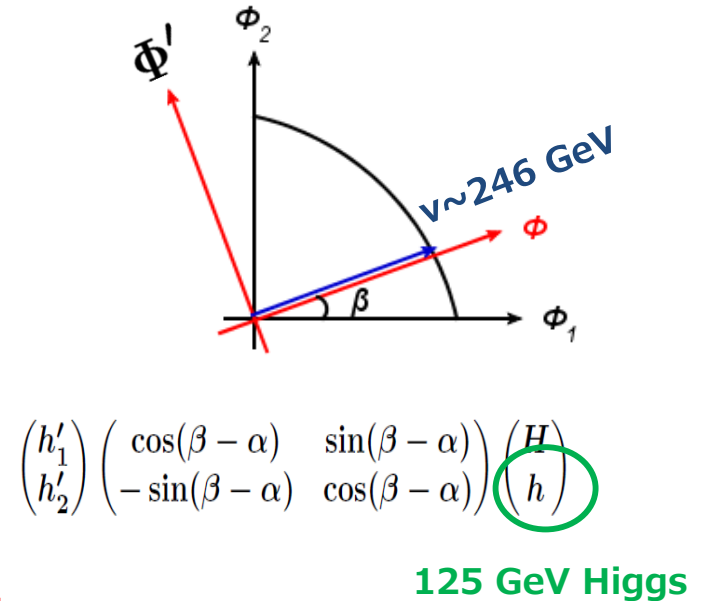
$\Phi = \begin{bmatrix} G^+ \\ \frac{1}{\sqrt{2}}(h'_1 + v + iG^0) \end{bmatrix}$

**CP-even Higgs**

**Charged Higgs**

$\Phi' = \begin{bmatrix} H^+ \\ \frac{1}{\sqrt{2}}(h'_2 + iA) \end{bmatrix}$

**CP-odd Higgs**



## □ Higgs boson masses

$$m_h^2 \sim \lambda v^2, \quad m_\Phi^2 \sim M^2 + \lambda' v^2 \quad (\Phi = H^\pm, A, H)$$

## □ Decoupling limit: $M^2 \rightarrow \infty$

## □ Alignment limit: $\sin(\beta - \alpha) \rightarrow 1$

# 2HDM with NFC

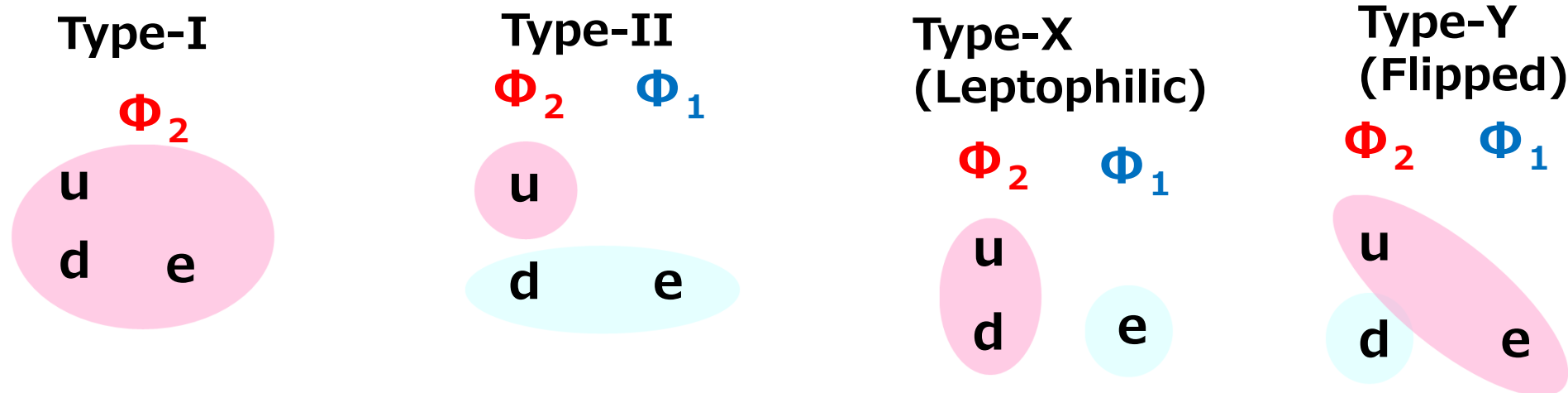
## □ Natural Flavor Conservation (NFC) Scenario

$\Phi_{u,d,e}$  : Either  $\Phi_1$  or  $\Phi_2$

$$-\mathcal{L}_Y = Y_u \bar{Q}_L (i\sigma_2) \Phi_u^* u_R + Y_d \bar{Q}_L \Phi_d d_R + Y_e \bar{L}_L \Phi_e e_R + \text{h.c.}$$

## □ This can be realized by imposing a (softly-broken) $Z_2$ symmetry.

*Barger, Hewett, Phillips, PRD41 (1990); Grossman, NPB426 (1994)*



# Yukawa couplings

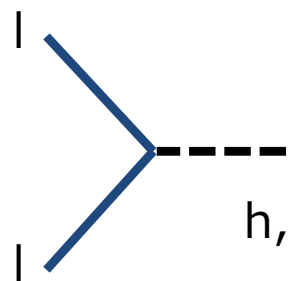
## Lepton Yukawa interactions

$$\mathcal{L}_{\text{lep}} = -Y_\ell \bar{L}_L \Phi_\ell \ell_R = -Y_\ell \bar{L}_L (\Phi + \xi_\ell \Phi') \ell_R$$

$$\frac{\sqrt{2}}{v} \begin{pmatrix} m_e & 0 & 0 \\ 0 & m_\mu & 0 \\ 0 & 0 & m_\tau \end{pmatrix}$$

$\cot\beta$  (Type-I, Y)

$-\tan\beta$  (Type-II, X)



$h, H, A$

$$\frac{m_\ell}{v} (s_{\beta-\alpha} + c_{\beta-\alpha} \xi_\ell)$$

(for  $h$ )

$$\frac{m_\ell}{v} (c_{\beta-\alpha} - s_{\beta-\alpha} \xi_\ell)$$

(for  $H$ )

$$i \frac{m_\ell}{v} \xi_\ell$$

(for  $A$ )

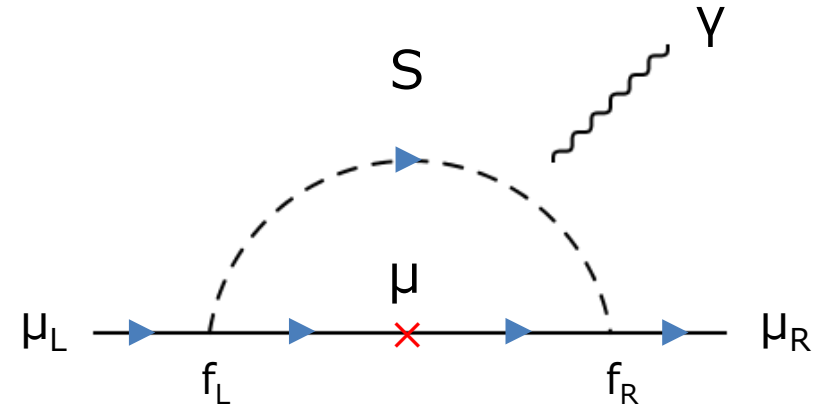
# 2HDMs with $Z_2$

▣ New interactions:

$$\mathcal{L}_{\text{new}} = \bar{F}(f_L P_L + f_R P_R)\mu S + \text{h.c.}$$

$\uparrow$   
Muon
 $\uparrow$   
 $h, H, A$  and  $H^\pm$

$\frac{m_\mu}{v}\xi_\ell$



This corresponds to the chirality **suppressed** case:

$$\sim \frac{1}{8\pi^2} \left( \frac{m_\mu}{v} \xi_\ell \right)^2 \times \left( \frac{m_\mu}{m_H} \right)^2 \sim 2 \times 10^{-9} \times \left( \frac{100 \text{ GeV}}{m_H} \right) \times \left( \frac{\xi_\ell}{1000} \right)^2$$

Type-II and Type-X 2HDMs:  $\xi_\ell \rightarrow \tan \beta$

However, such a super large  $\tan \beta$  is not allowed in these 2HDMs (e.g., LHC, Flavor exp. ).

# 2HDMs with $Z_2$

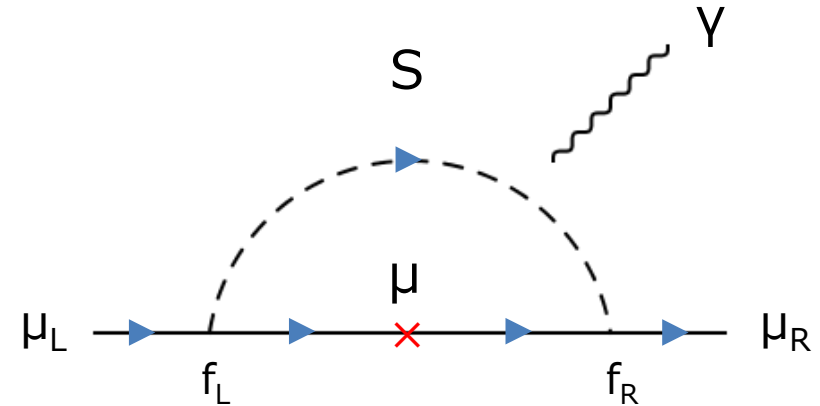
▣ New interactions:

$$\mathcal{L}_{\text{new}} = \bar{F}(f_L P_L + f_R P_R)\mu S + \text{h.c.}$$

Muon
h, H, A and  $H^\pm$

$\frac{m_\mu}{v}\xi_\ell$

*(Note: Red dotted arrows point from the  $\frac{m_\mu}{v}\xi_\ell$  term to the  $f_L$  and  $f_R$  vertices in the diagram below.)*



This corresponds to the chirality **suppressed** case:

$$\sim \frac{1}{8\pi^2} \left( \frac{m_\mu}{v} \xi_\ell \right)^2 \times \left( \frac{m_\mu}{m_H} \right)^2 \sim 2 \times 10^{-9} \times \left( \frac{100 \text{ GeV}}{m_H} \right) \times \left( \frac{\xi_\ell}{1000} \right)^2$$

Type-II and Type-X 2HDMs:  $\xi_\ell \rightarrow \tan \beta$

However, such a super large  $\tan \beta$  is not allowed in these 2HDMs (e.g., LHC, Flavor exp. ).

*“Muon Specific 2HDM” Abe, Sato, KY, 1705.01469 (JHEP)*

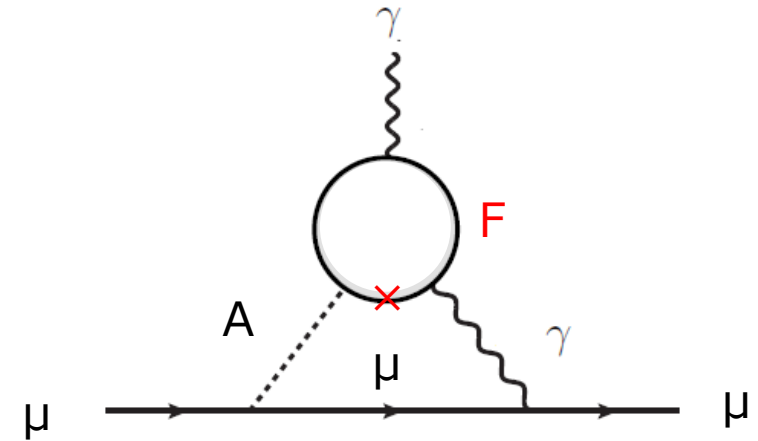
- Way out 1: Extending  $Z_2$  and having  $\tan \beta$  enhancement only in the muon Yukawa coupling.
- Way out 2: 2-loop Barr-Zee contribution in the Type-X 2HDM.
- Way out 3: Forget about the  $Z_2$  symmetry.



# 2HDMs with $Z_2$

## 2-loop Barr-Zee contribution

$$a_\mu \sim \left( \frac{1}{16\pi^2} \right)^2 16e^2 \frac{m_\mu^2}{v^2} \frac{m_F^2}{m_A^2} \xi_\ell \xi_F$$



## In the Type-X case, tau-loop ( $F = \tau$ ) can be important.

$$\text{c.f. 1-loop contribution} \quad \sim \frac{1}{16\pi^2} \times \frac{m_\mu^2}{v^2} \times \frac{m_\mu^2}{m_H^2} \times (\tan \beta)^2$$

*Chun et.al., 1409.3199*

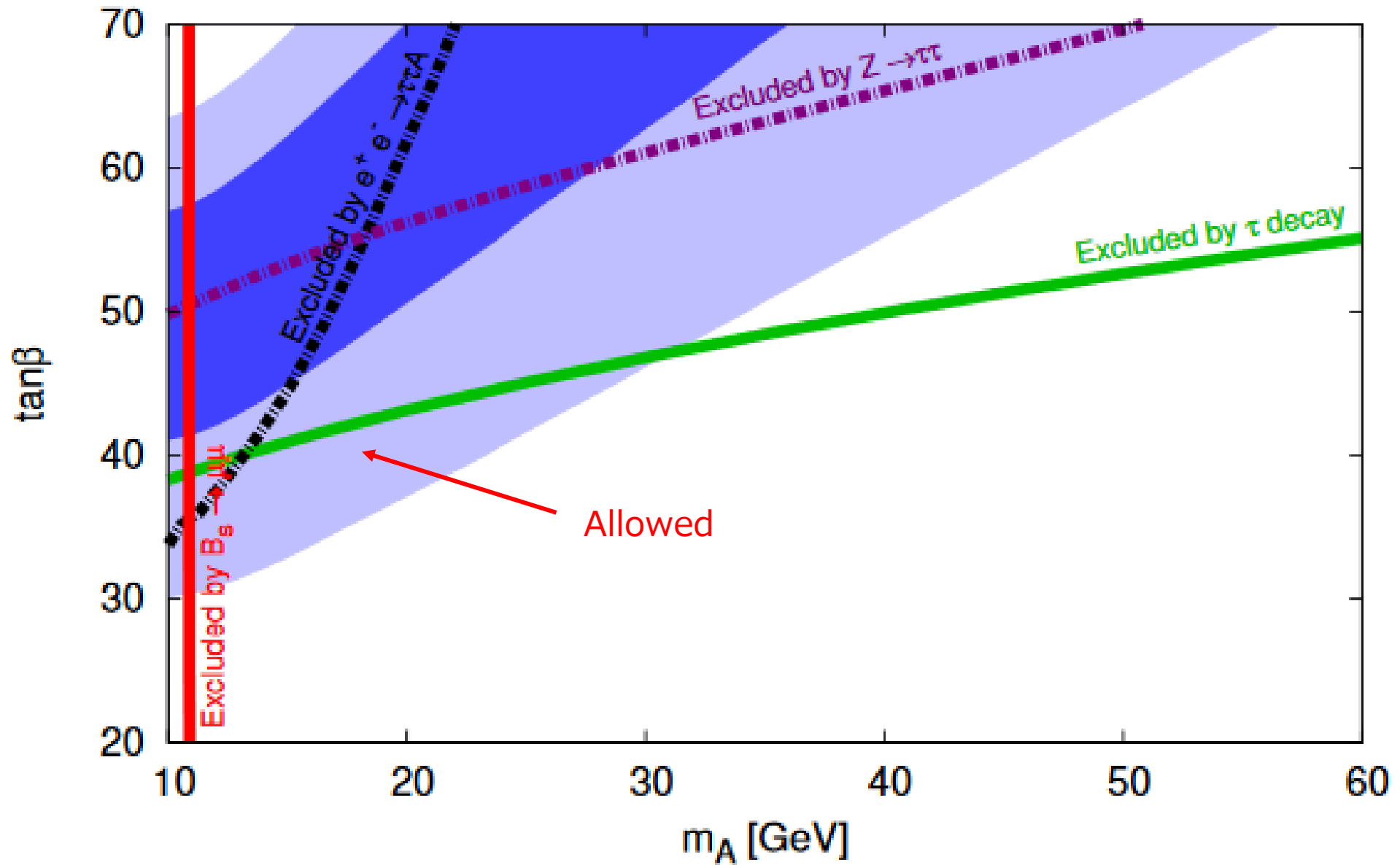
*Abe, Sato, KY, 1504.07059*

$$a_\mu \sim \left( \frac{1}{16\pi^2} \right)^2 16e^2 \frac{m_\mu^2}{v^2} \frac{m_\tau^2}{m_A^2} (\tan \beta)^2 \sim 10^{-9} \times \left( \frac{\tan \beta}{50} \right)^2 \times \left( \frac{20 \text{ GeV}}{m_A} \right)^2$$

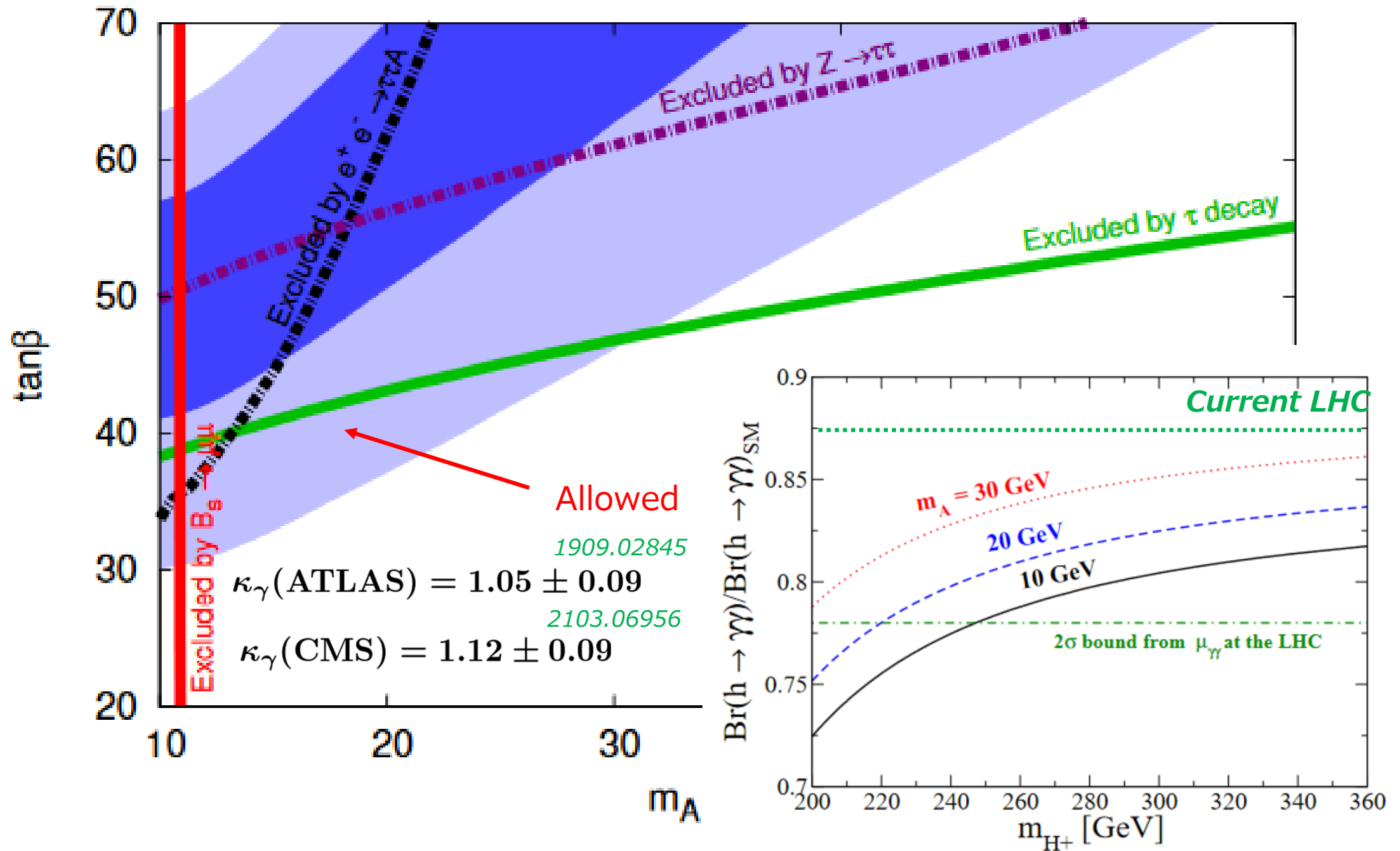
*Chun, Kang, Takeuchi, Tsai, 1507.08067*

Light CP-odd Higgs boson is required, which mainly decays into a tau pair.

$$m_{H^0} = m_{H^\pm} = 300 \text{ GeV}$$



$$m_{H^0} = m_{H^\pm} = 300 \text{ GeV}$$



# 2HDMs without $Z_2$

Omura, Senaha, Tobe, 1502.07824

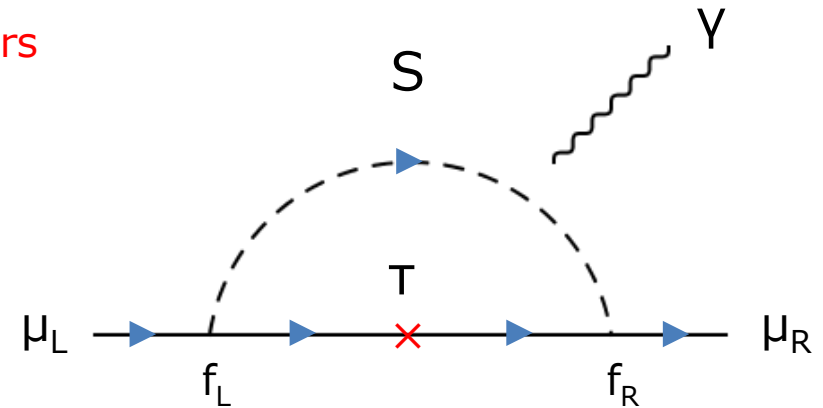
## □ New interactions:

$$\mathcal{L}_{\text{new}} = \bar{F}(f_L P_L + f_R P_R)\mu S + \text{h.c.}$$

$T$

$h, H, A$  and  $H^\pm$

Free parameters



$$a_\mu^{\text{NP}} \sim \frac{f_L f_R}{8\pi^2} \left( \frac{m_\mu}{M_F} \right) \left( \frac{M_F}{M_S} \right)^2$$

$$\sim \frac{1}{16\pi^2} \times \frac{m_\mu}{m_\tau} \times \frac{m_\tau^2}{m_H^2} \times (f_L f_R) \sim 10^{-9} \times \left( \frac{v}{m_H} \right) \times \left( \frac{f_L f_R}{0.3^2} \right)$$

• We can explain  $(g-2)_\mu$ , but the flavor problem may happen ...

• Can we explain  $(g-2)_\mu$  and **DM** at the same time?

**Yes!**

# Contents

I. Introduction

II. Muon  $g-2$  in BSMs

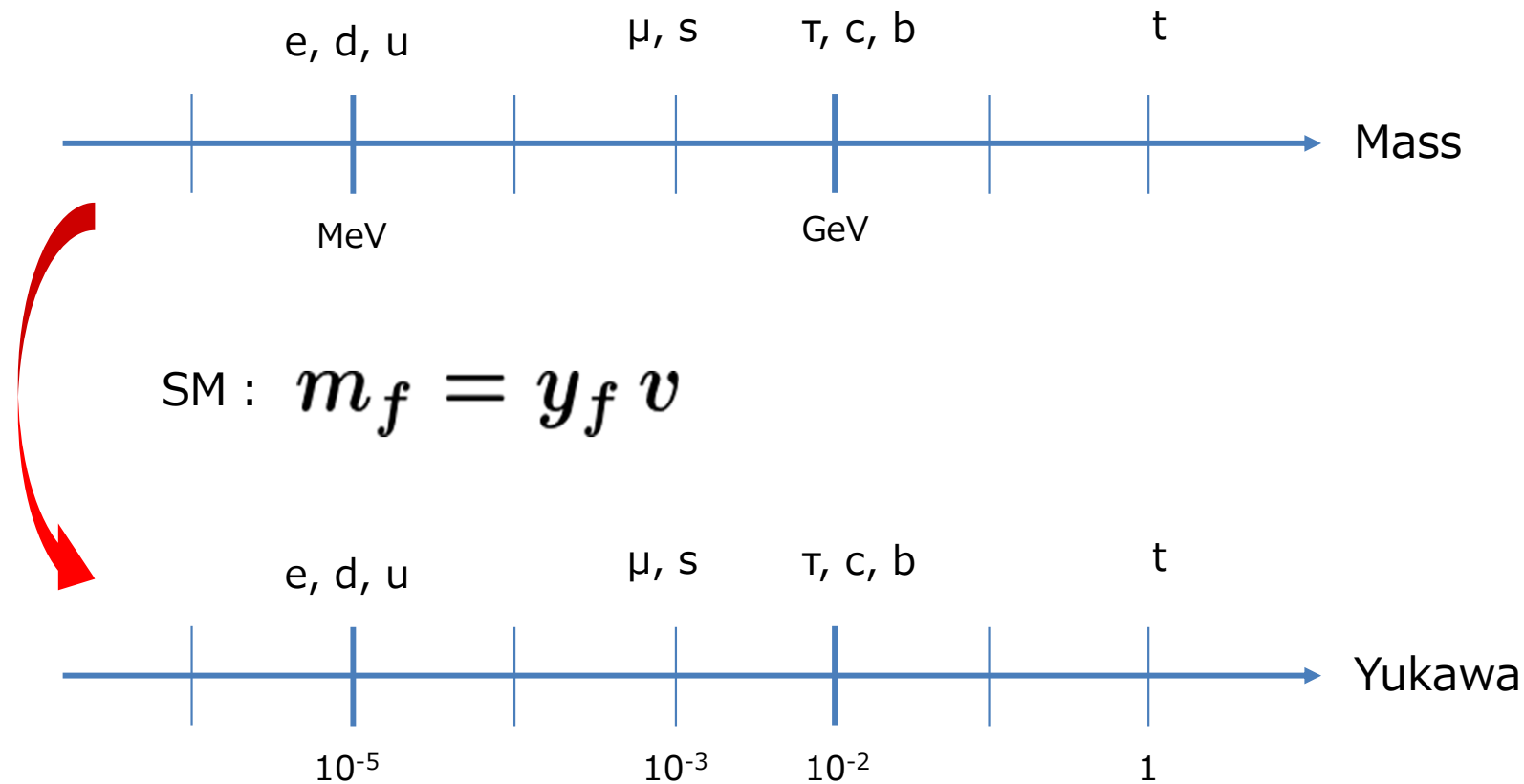
III. Radiative Charged Seesaw Mechanism

*Cheng-Wei Chiang, KY, 2104.00890 [hep-ph] (PRD)*

IV. Collider phenomenology

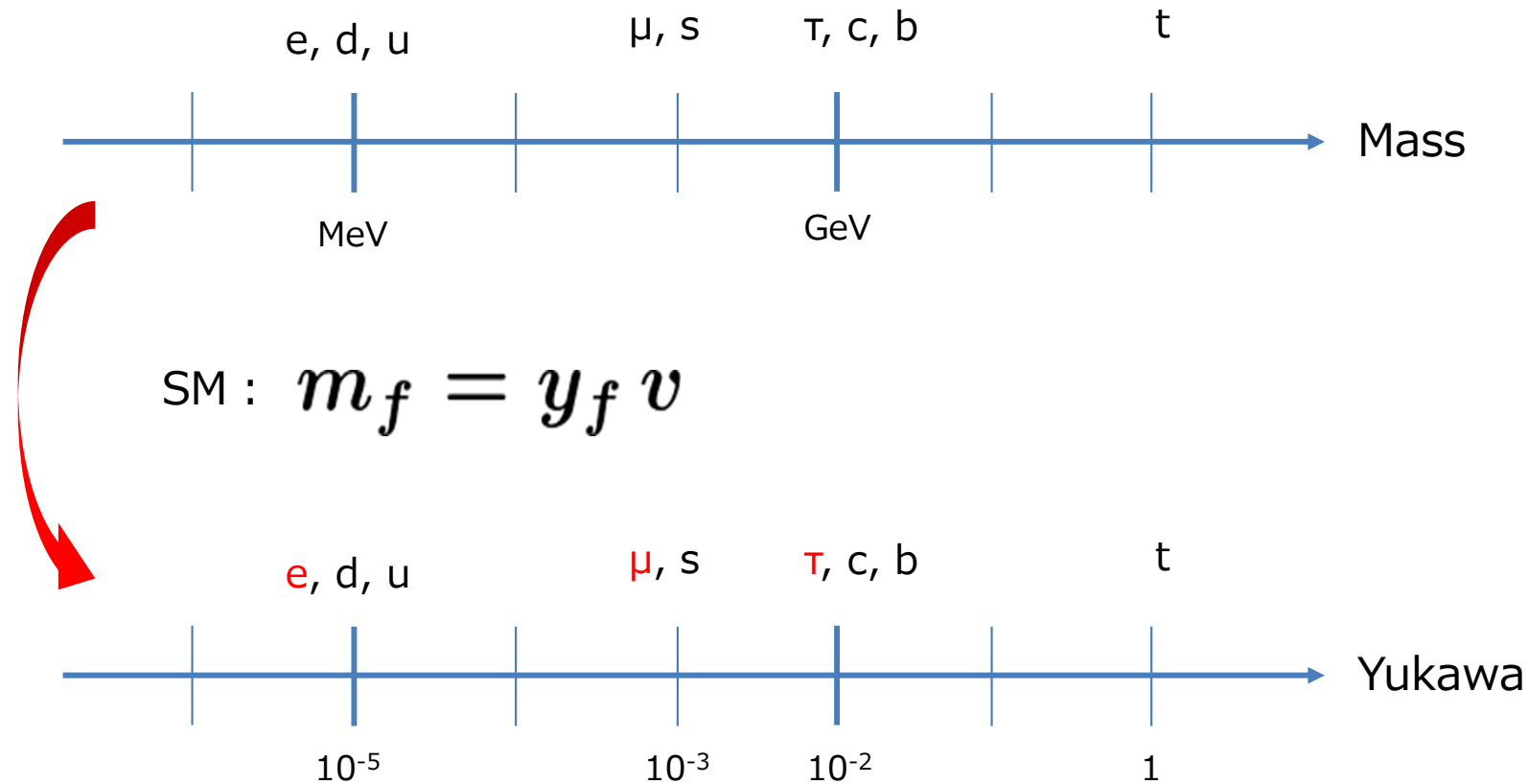
V. Summary

# Fermion Mass Hierarchy



Can we explain the mass hierarchy with  $O(1)$  couplings?

# Fermion Mass Hierarchy

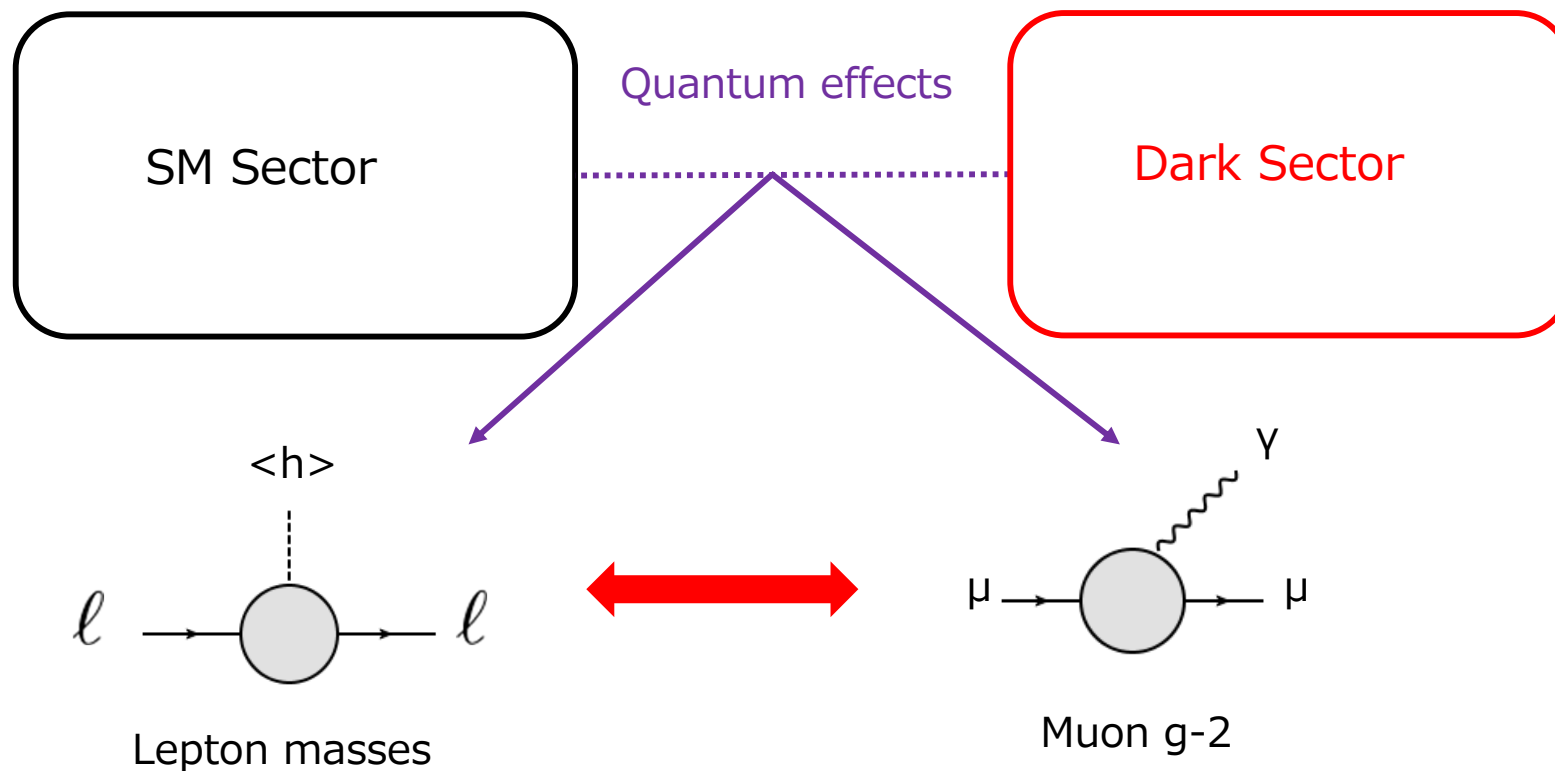


Can we explain the mass hierarchy with  $O(1)$  couplings?

In particular, we focus on the **charged lepton mass** hierarchy.

# Dark sector as origin of lepton masses

- Introduction of “dark sector” to the SM



$$\sim \left( \frac{1}{16\pi^2} \right)^N \times \left( \frac{v}{M} \right) \times Y_{\text{New}} \quad \text{“Radiative Charged Seesaw”}$$

Dark sector simultaneously explains tiny lepton masses and  $(g-2)_\mu$  anomaly.



# How to realize the mechanism?

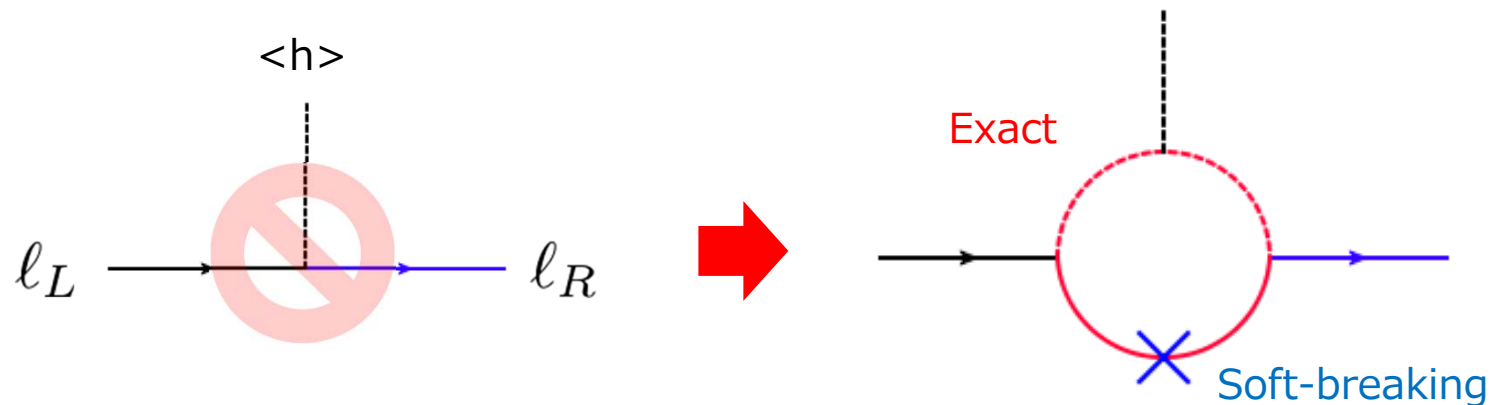
1. Dark sector can be defined by introducing a  $Z_2$  symmetry.

$$\Psi_{\text{SM}} \rightarrow +\Psi_{\text{SM}}, \quad \Psi_{\text{Dark}} \rightarrow -\Psi_{\text{Dark}}$$

2. Tree level Yukawa should be forbidden by another  $Z_2'$  symmetry which has to be softly-broken to generate a finite mass.

3. Chirality flip should be introduced in the loop.

Vector-like mass  
Higgs mechanism



# Model

- We consider the case with the tree level tau mass and 1-loop induced  $\mu/e$  masses.

	Fermions				Scalars		
Fields	$(L_L^e, L_L^\mu, L_L^\tau)$	$(e_R, \mu_R, \tau_R)$	$(F_L^e, F_L^\mu)$	$(F_R^e, F_R^\mu)$	$H$	$\Phi_L$	$\Phi_R$
$U(1)_\ell$ (exact)	$(q_e, q_\mu, 0)$	$(q_e, q_\mu, 0)$	$(q_e, q_\mu)$	$(q_e, q_\mu)$	0	0	0
$Z_2$ (exact)	$(+, +, +)$	$(+, +, +)$	$(-, -)$	$(-, -)$	+	-	-
$Z'_2$ (soft-breaking)	$(+, +, +)$	$(-, -, +)$	$(+, +)$	$(-, -)$	+	-	-

Forbid LFVs

Forbid tree level mixing

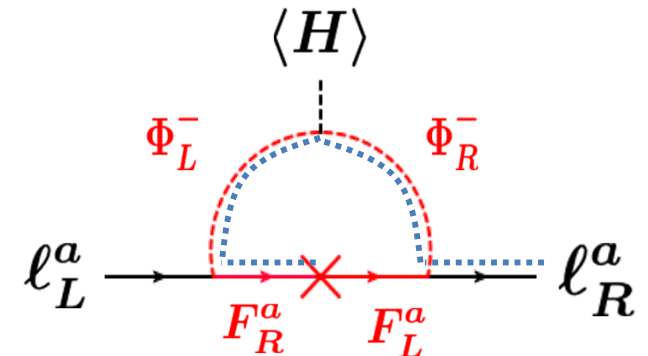
Forbid tree level Yukawa

Stabilize the DM candidate

New particles (Dark sector)

- Lagrangian for the lepton sector

$$\mathcal{L} = -y_\tau \bar{L}_L^\tau H \tau_R - \sum_{a=e,\mu} (M^a \bar{F}_L^a F_R^a + f_L^a \bar{L}_L^a \Phi_L F_R^a + f_R^a \bar{\ell}_R^a \Phi_R F_L^a) - \mu H \Phi_L \Phi_R$$



# Charge assignment under $SU(2)_I \times U(1)_Y$

- We list the possible sets of the  $SU(2)_I \times U(1)_Y$  charge.

$$F \sim (I_F, Y_F) \quad \Phi_{L,R} \sim (I_{L,R}, Y_{L,R}) \quad \bar{L}_L \Phi_L F_R$$

$$\text{with } Y_L = -1/2 - Y_F, \quad Y_R = -1 - Y_F \quad \bar{\ell}_R \Phi_R F_L$$

$(I_F, Y_F)$	$(I_L, Y_L)$	$(I_R, Y_R)$	Sign of $\Delta a_\ell$
<b>(1, 0)</b>	<b>(2, -1/2)</b>	<b>(1, -1)</b>	+
<b>(1, -1)</b>	<b>(2, 1/2)</b>	<b>(1, 0)</b>	-
<b>(2, 1/2)</b>	<b>(1 or 3, -1)</b>	<b>(2, -3/2)</b>	+
<b>(2, -1/2)</b>	<b>(1 or 3, 0)</b>	<b>(2, -1/2)</b>	- or $\pm$
<b>(2, -3/2)</b>	<b>(1 or 3, 1)</b>	<b>(2, 1/2)</b>	-
<b>(3, 1)</b>	<b>(2, -3/2)</b>	<b>(3, -2)</b>	+
<b>(3, 0)</b>	<b>(2, -1/2)</b>	<b>(3, -1)</b>	$\pm$
<b>(3, -1)</b>	<b>(2, 1/2)</b>	<b>(3, 0)</b>	-
<b>(3, -2)</b>	<b>(2, 3/2)</b>	<b>(3, 1)</b>	-

*E. Ma, 1311.3213*

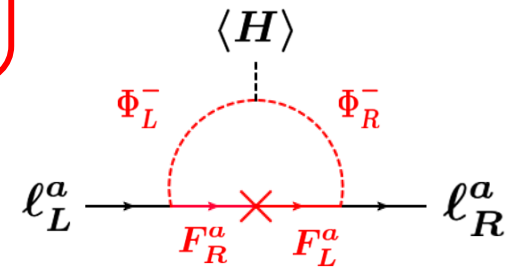
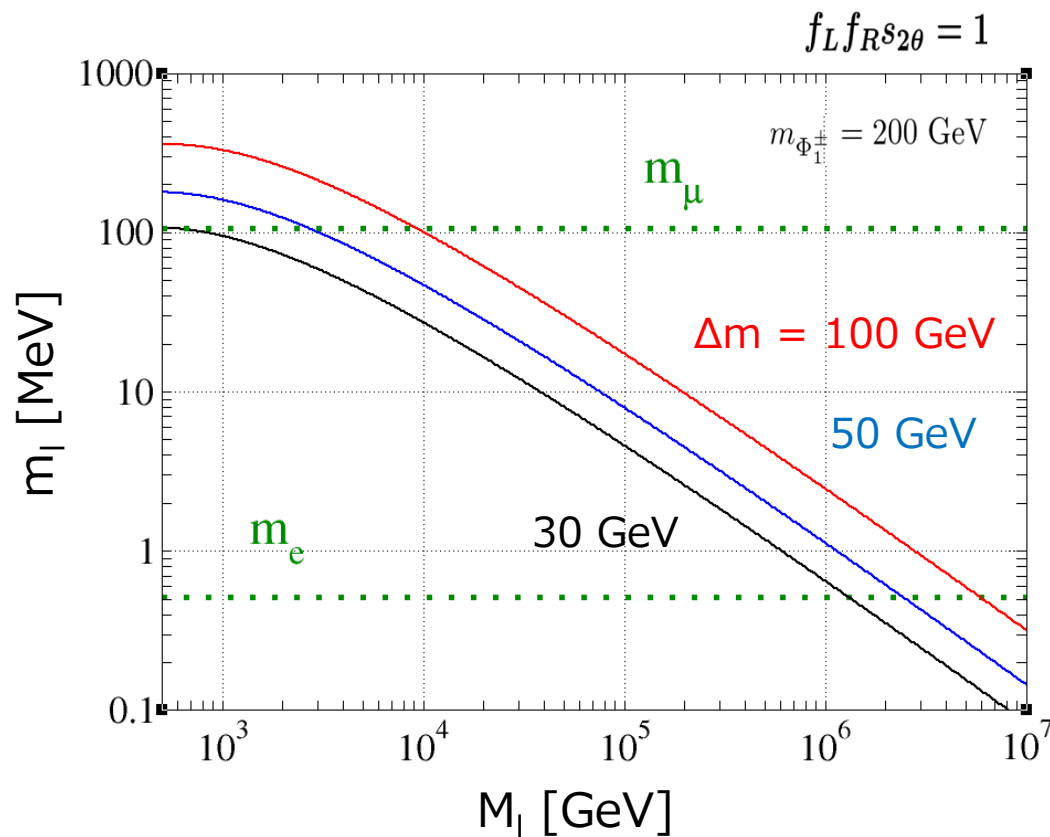
Common dark sector contributes to mass  
and  $(g-2)_\mu \rightarrow$  **Sign of  $(g-2)_\mu$  is fixed.**

Let us focus on the simplest case with  $F \sim (\mathbf{1}, 0)$ .

# Charged lepton masses

□ Mass eigenstates of the charged scalars  $\begin{pmatrix} \Phi_L^\pm \\ \Phi_R^\pm \end{pmatrix} = \begin{pmatrix} \cos \theta & -\sin \theta \\ \sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} \Phi_1^\pm \\ \Phi_2^\pm \end{pmatrix}$

$$m_\ell \simeq -\frac{f_L^\ell f_R^\ell s_{2\theta}}{16\pi^2} M_\ell \left[ \left( \frac{m_{\Phi_1^\pm}}{M_\ell} \right)^2 \ln \left( \frac{m_{\Phi_1^\pm}}{M_\ell} \right)^2 - \left( \frac{m_{\Phi_2^\pm}}{M_\ell} \right)^2 \ln \left( \frac{m_{\Phi_2^\pm}}{M_\ell} \right)^2 \right] \text{ for } M_\ell \gg m_{\Phi^\pm}$$



• Case with an O(1) coupling :

Muon mass  $\rightarrow M \sim \text{O}(1) \text{ TeV}$

Electron mass  $\rightarrow M \sim \text{O}(1) \text{ PeV}$

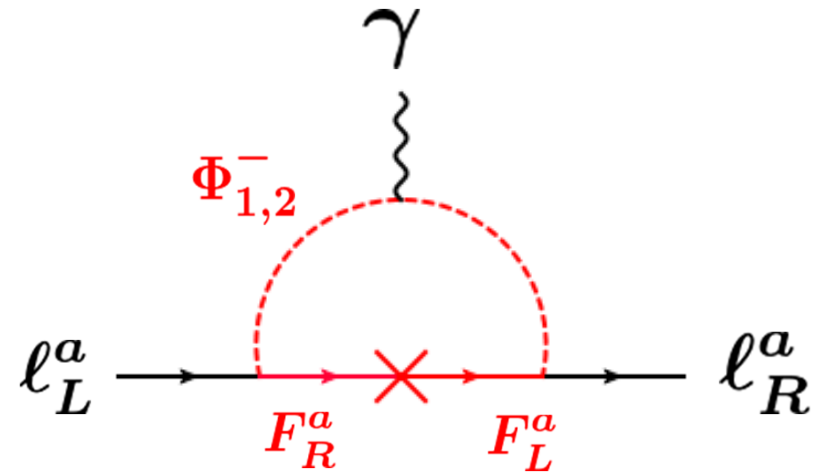
# Anomalous magnetic moments

$$\Delta a_\ell \equiv a_\ell^{\text{exp}} - a_\ell^{\text{SM}}$$

$$\Delta a_\mu = -\frac{f_L f_R s_{2\theta}}{16\pi^2} \frac{m_\ell}{M_\ell} \left[ H_S \left( \frac{m_{\Phi_1^\pm}^2}{M_\ell^2} \right) - H_S \left( \frac{m_{\Phi_2^\pm}^2}{M_\ell^2} \right) \right]$$

$$\simeq \left( \frac{m_\ell}{M_\ell} \right)^2 \left[ 1 + \frac{3}{2} \frac{1}{1 + \ln(m_\Phi^2/M_\ell^2)} \right]$$

for  $M_l \gg m_{\Phi_1} = m_{\Phi_2} (=m_\Phi)$



- It gives a **positive** contribution to  $(g-2)_l$ .
- The dependence of the new Yukawa couplings does not explicitly appear.

# Anomalous magnetic moments

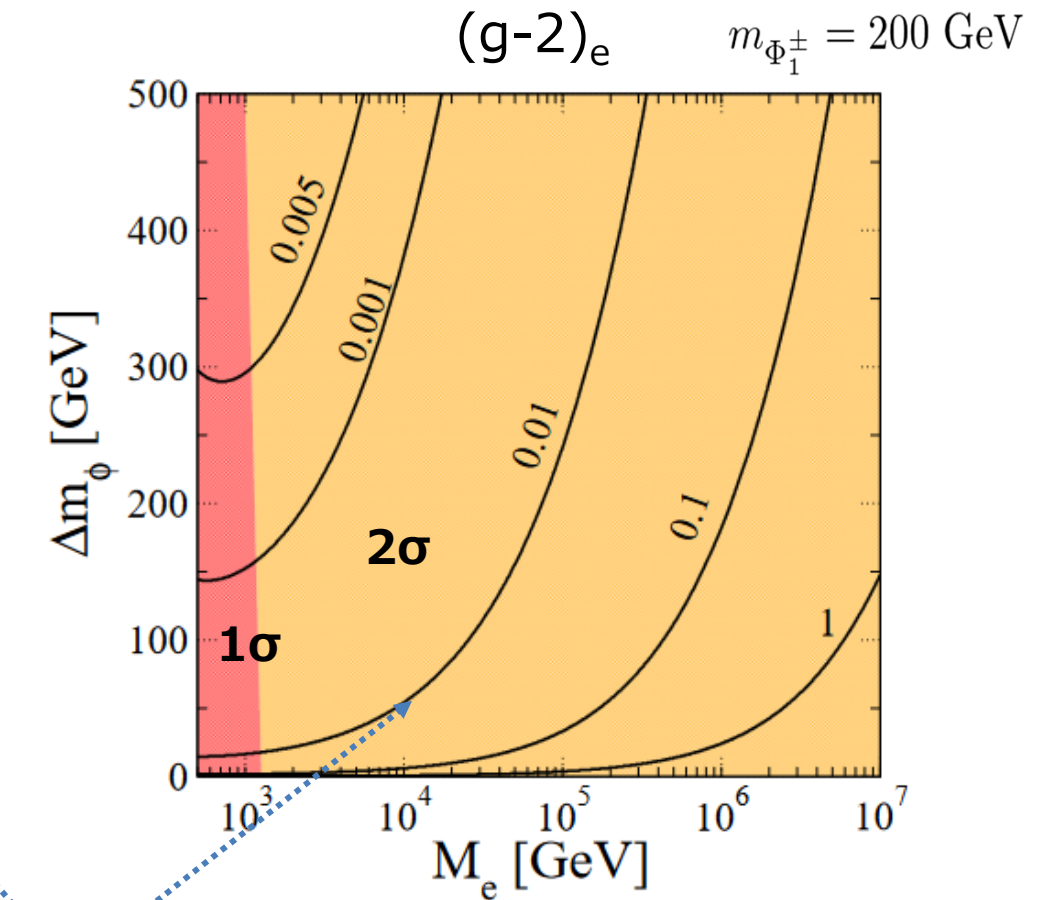
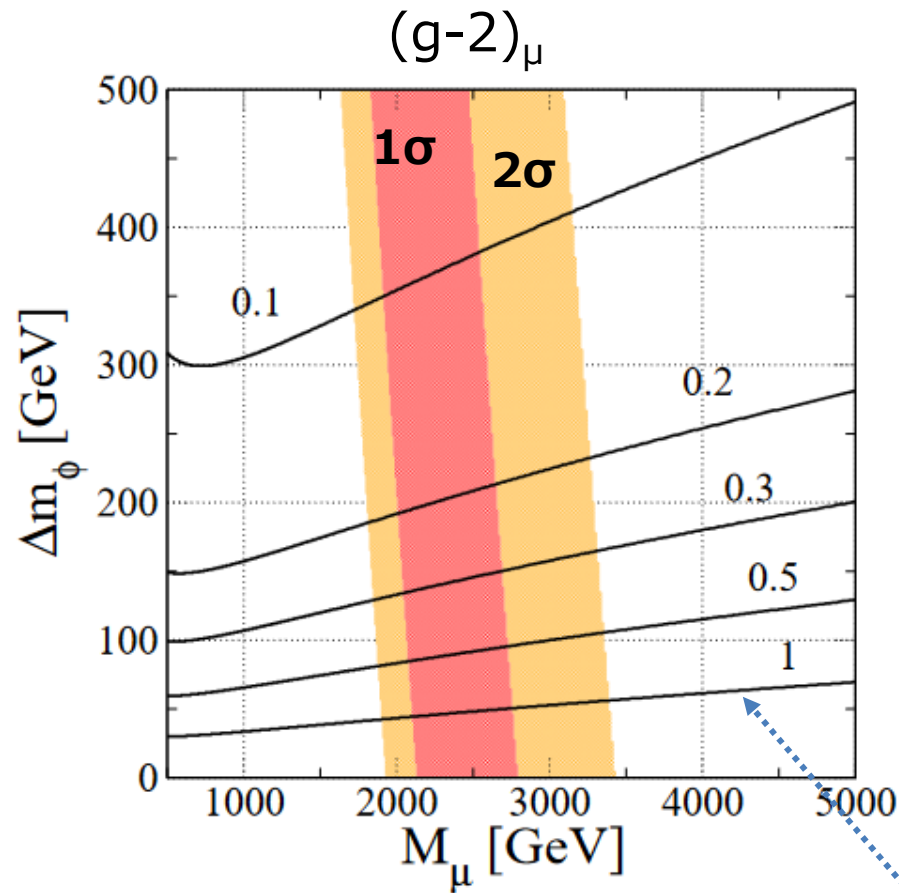
Cheng-Wei Chiang, KY, 2104.00890 [hep-ph]

FNAL (2021)

$$\Delta a_\mu = (251 \pm 59) \times 10^{-11} \quad (4.2\sigma)$$

Nature, 588 (2020)

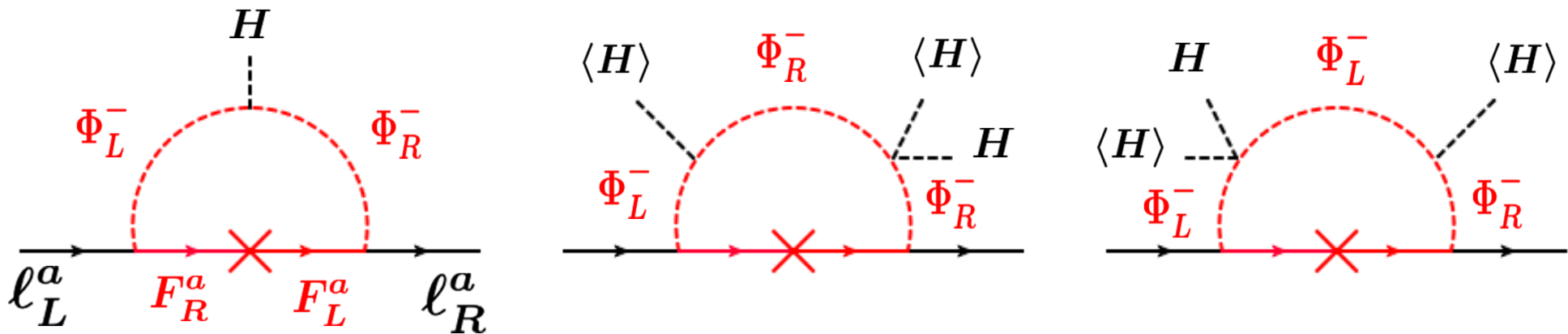
$$\Delta a_e = (4.8 \pm 3.0) \times 10^{-13} \quad (1.6\sigma)$$



$f_L f_R s_{2\theta}$

# Yukawa coupling

- Yukawa coupling does not simply obey  $y_f = m_f/v$ .



$$y_\ell \simeq \frac{m_\ell}{v} \underbrace{\frac{1}{|1 + \ln(m_{\Phi_1}^2/M_\ell^2)|} \left[ 2 + \ln \left( \frac{m_{\Phi_1}^2}{M_\ell^2} \right) \right] + \frac{v^2 \lambda_0}{m_{\Phi_1^\pm}^2}}_{\equiv \kappa_\ell} \quad \text{for } M_1 \gg m_{\Phi_1} = m_{\Phi_2} (=m_\Phi)$$

$$\lambda_0 = \frac{m_{\text{DM}}^2}{v^2} + \lambda_{\text{DM}} - \frac{\lambda_{\text{HR}}}{2}$$

$$V = \lambda_{\text{HR}}(H^\dagger H)(\Phi_R^* \Phi_R) + \dots$$

- The deviation is not suppressed by the loop factor, and it can be sizable.

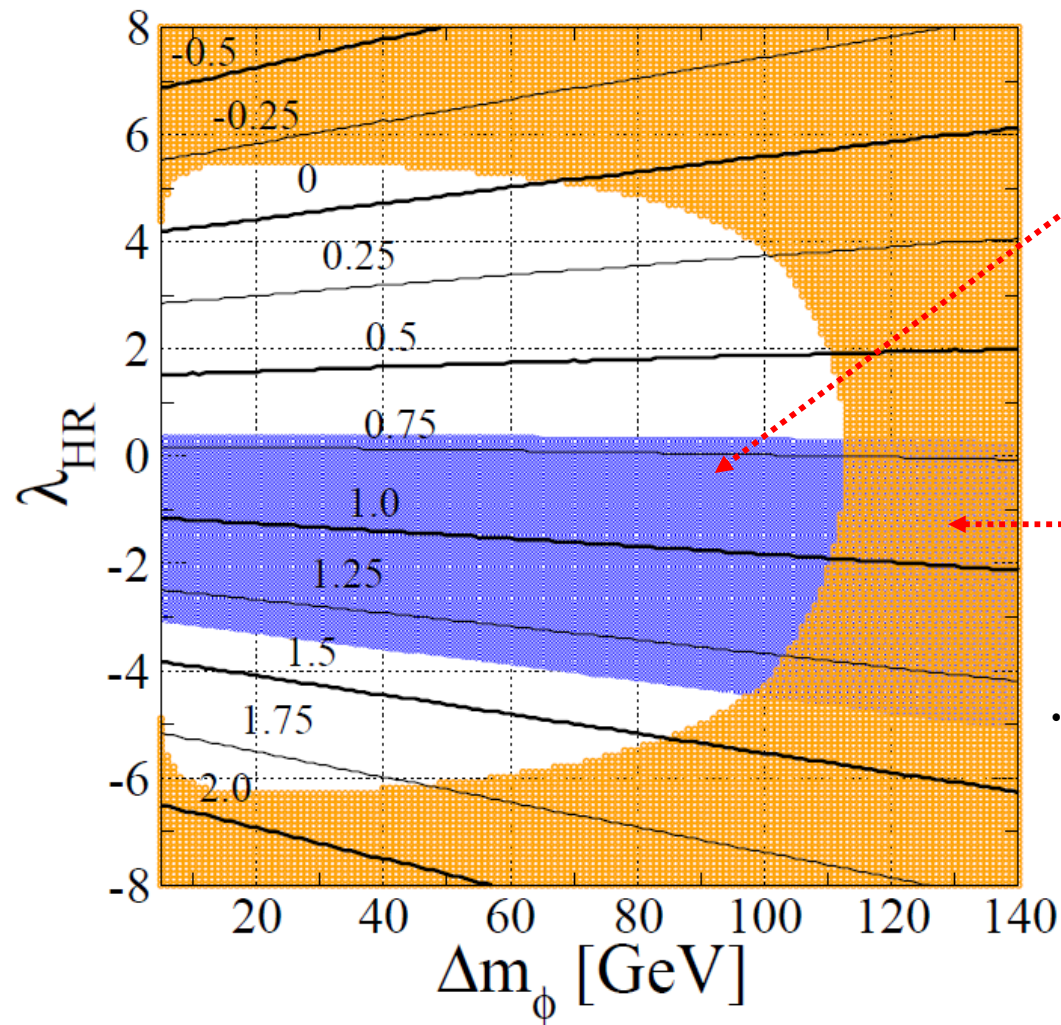


# Yukawa coupling

Cheng-Wei Chiang, KY, 2104.00890 [hep-ph]

▣ Contours for  $\kappa_\ell = y_\ell / y_\ell^{\text{SM}}$

$$m_{\Phi_1^\pm} = 200 \text{ GeV}$$



Allowed (95%CL) by the signal strength  
for  $pp \rightarrow h \rightarrow \mu\mu$ .

ATLAS, 2007.07830 [hep-ex]

CMS-PAS-HIG-19-006

Excluded by perturbativity bound  
by using 1-loop RGEs.

• Future measurements of the  $\mu$ -Yukawa:

→ HL-LHC  $\sim 7\%$ , ILC(250)  $\sim 5\%$

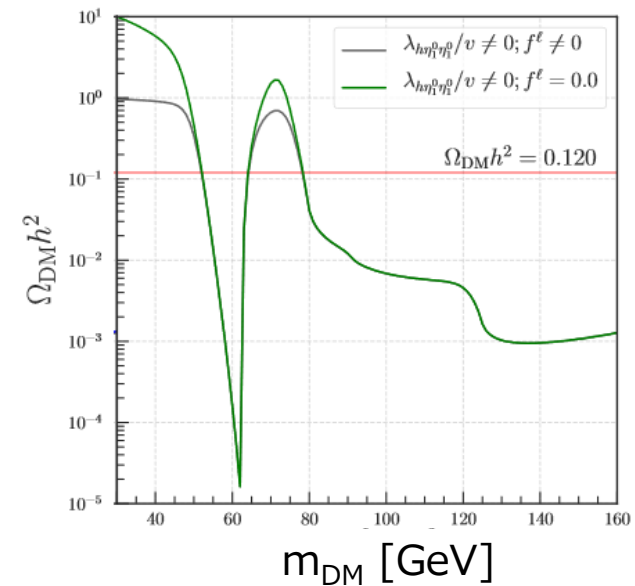
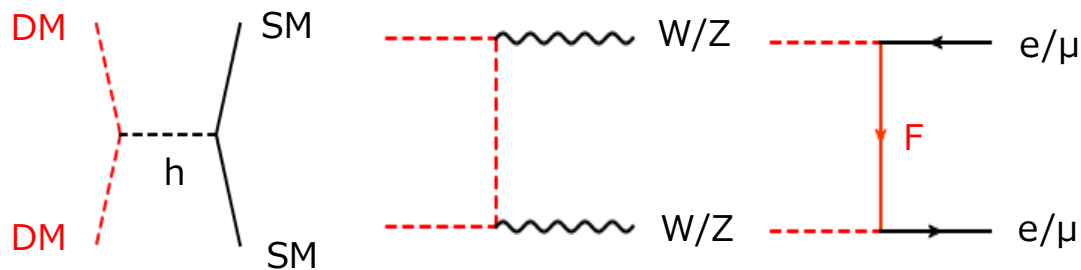


# Dark matter

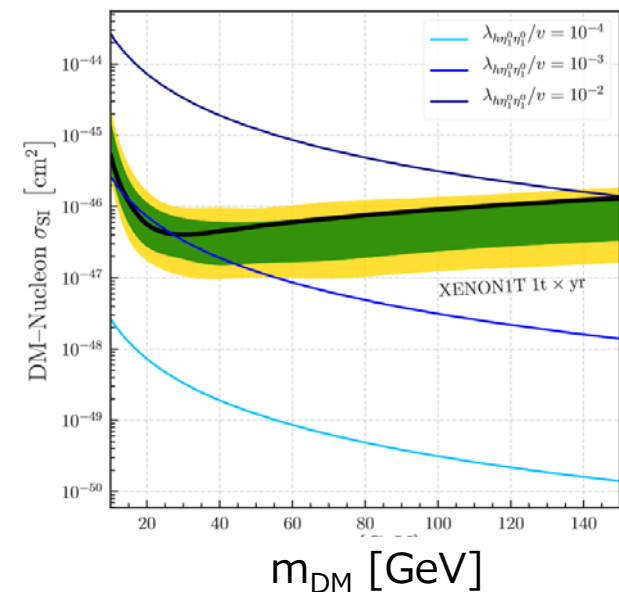
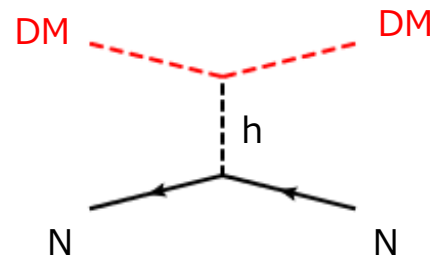
Kai-Feng Chen, Cheng-Wei Chiang, KY, 2006.07929 (JHEP)

□  $DM = \text{Re}(\Phi_L^0)$

□ Annihilation processes



□ Direct searches



- There are solutions at  $m_{DM} \sim 63$  GeV and 80 GeV.
- We need  $|\lambda_{DM}| < 10^{-3}$  to avoid direct search constraint.

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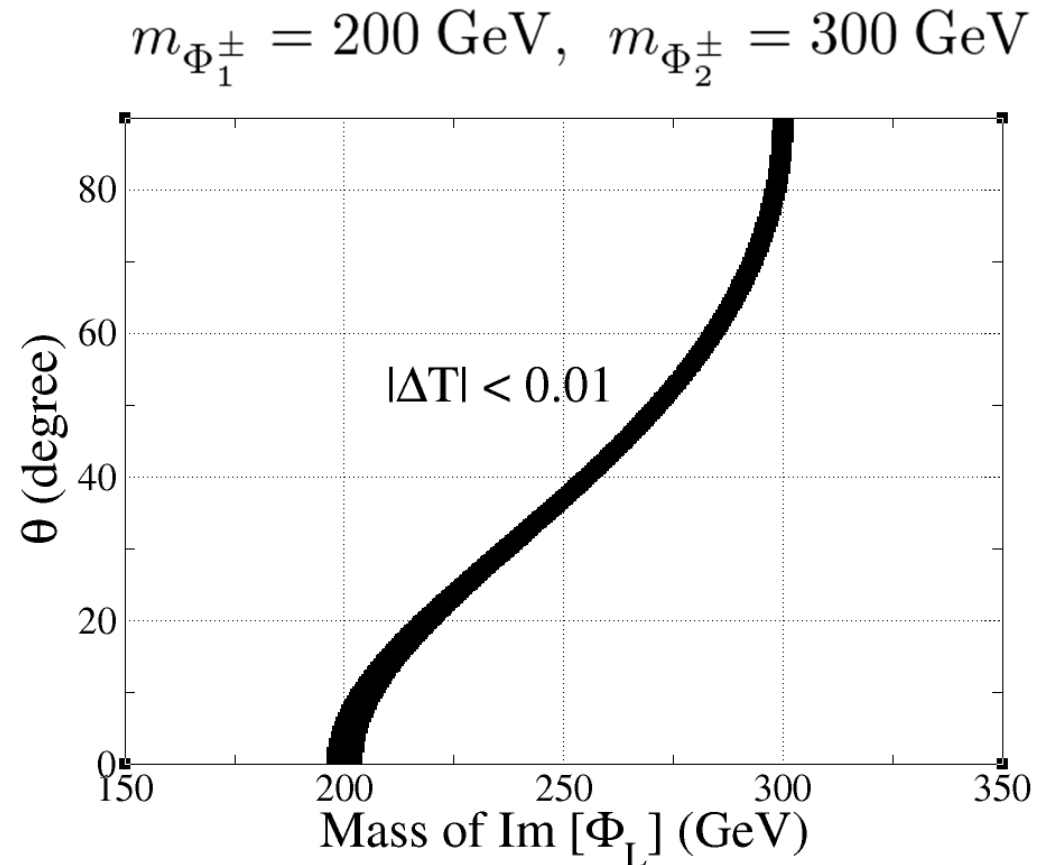
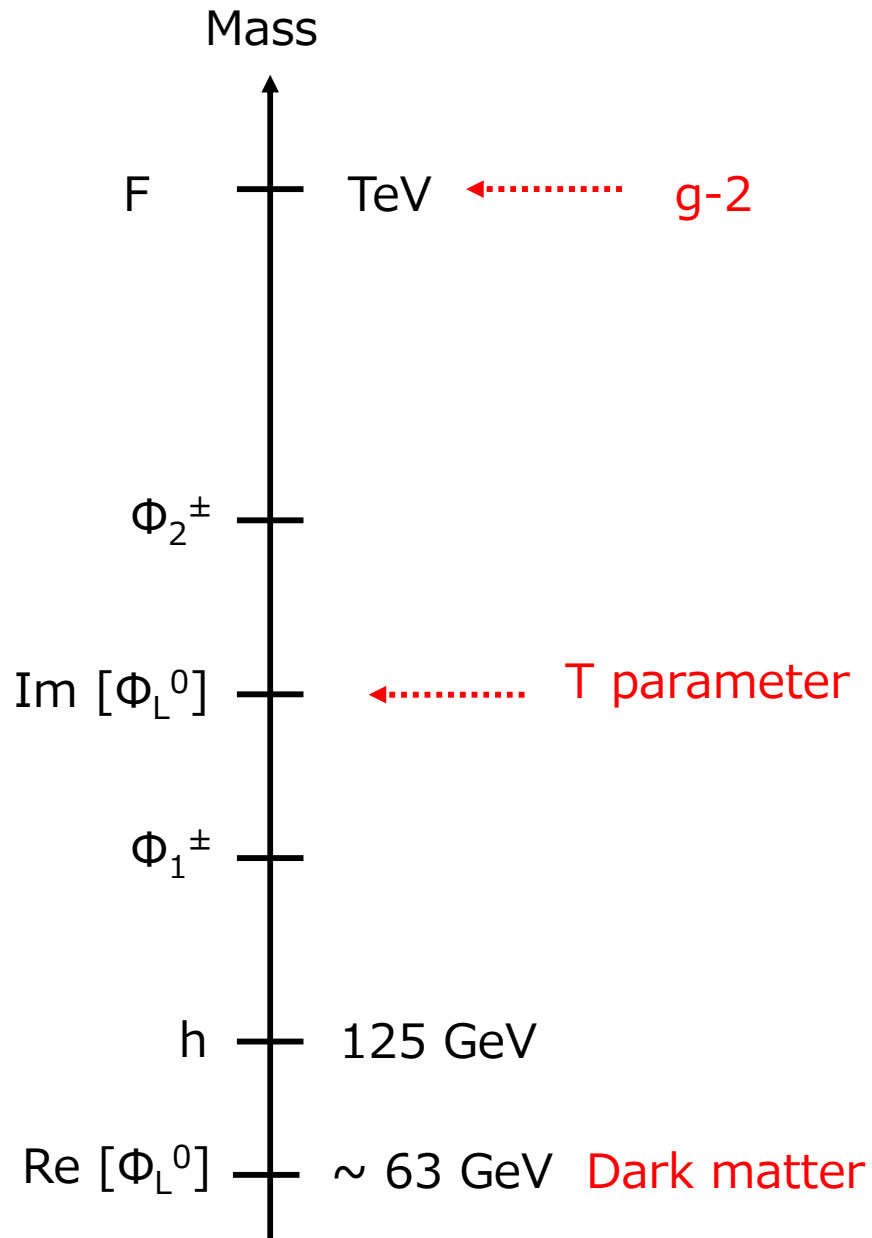
III. Radiative Charged Seesaw Mechanism

IV. Collider phenomenology (on going)

*Cheng-Wei Chiang, Ryomei Obuchi, KY, work in progress*

V. Summary

# Mass Spectrum

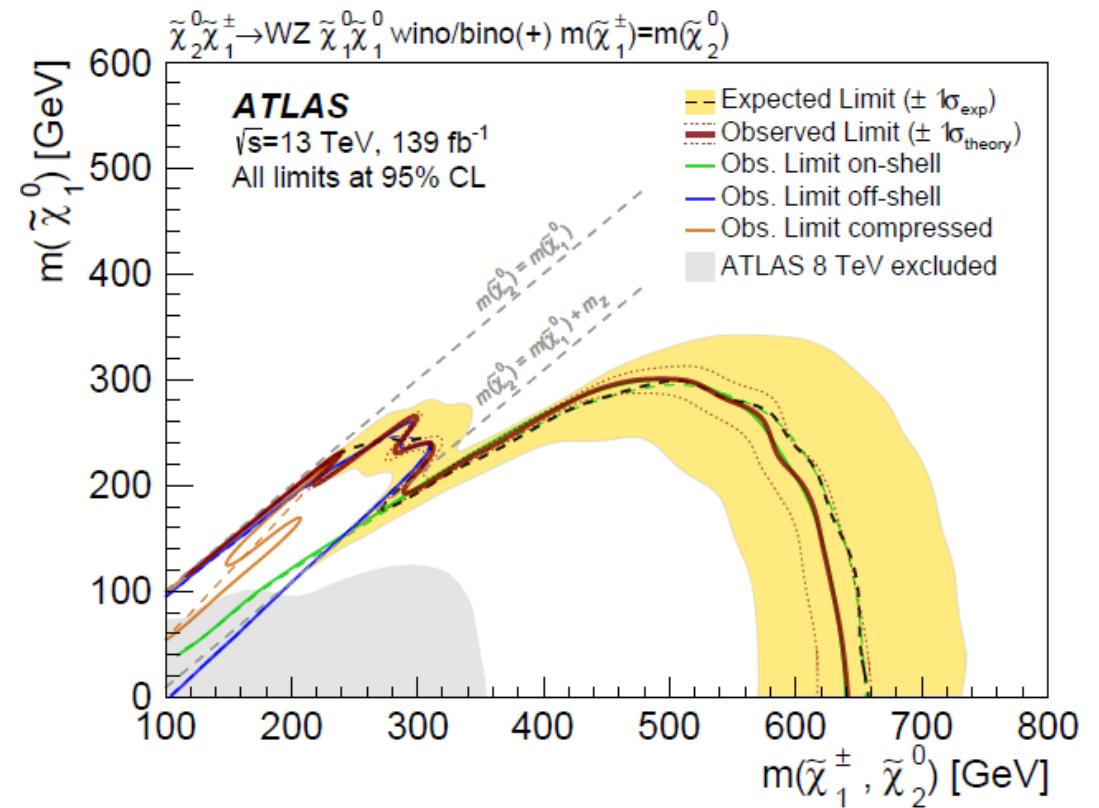
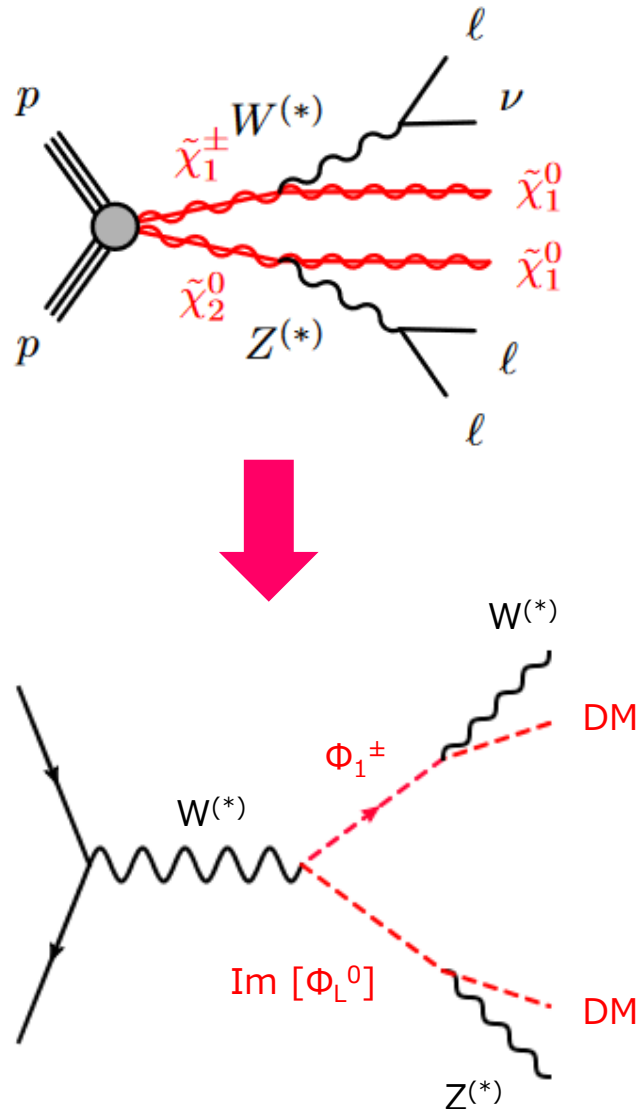


- Dark scalars typically decay into DM + W/Z/h.

# LHC Search

2106.01676 (ATLAS)

## □ SUSY search

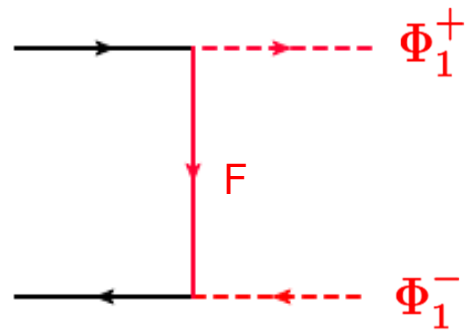
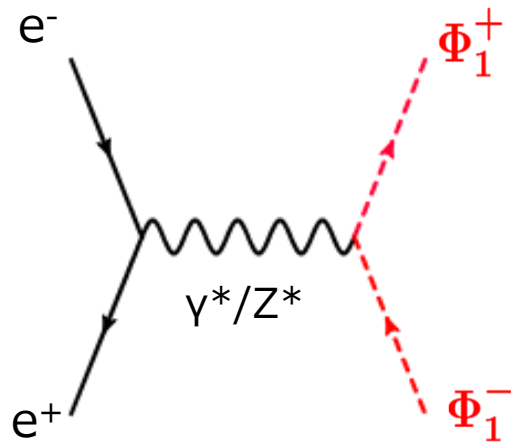


- Large region has already been excluded, but a careful study is needed to apply the limit to our model.
- If  $\Phi_1^\pm$  are singlet-like, we may be able to avoid the bound even for small mass region.

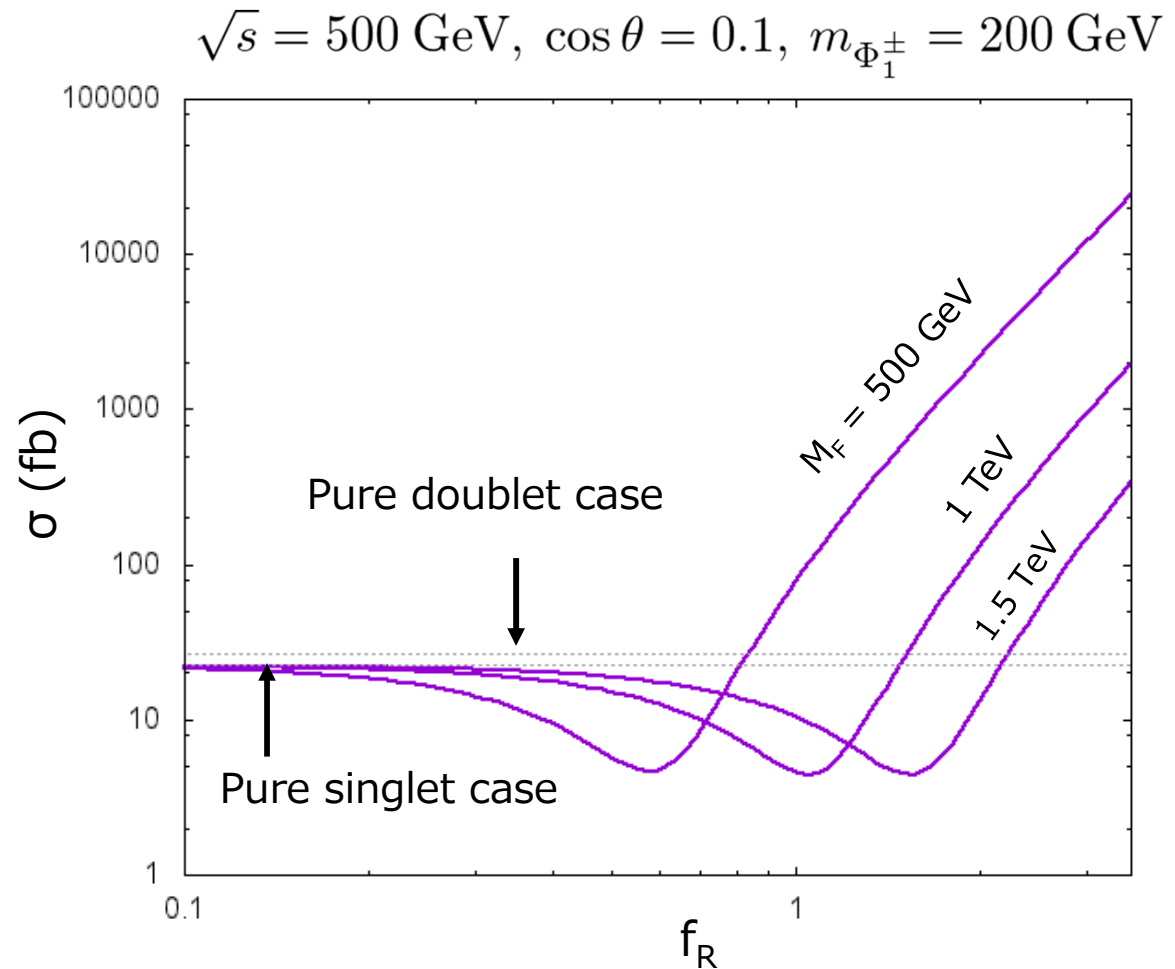
# ILC Search

*Cheng-Wei Chiang, Ryomei Obuchi, KY, work in progress*

- Lighter charged scalars can be produced in pair.



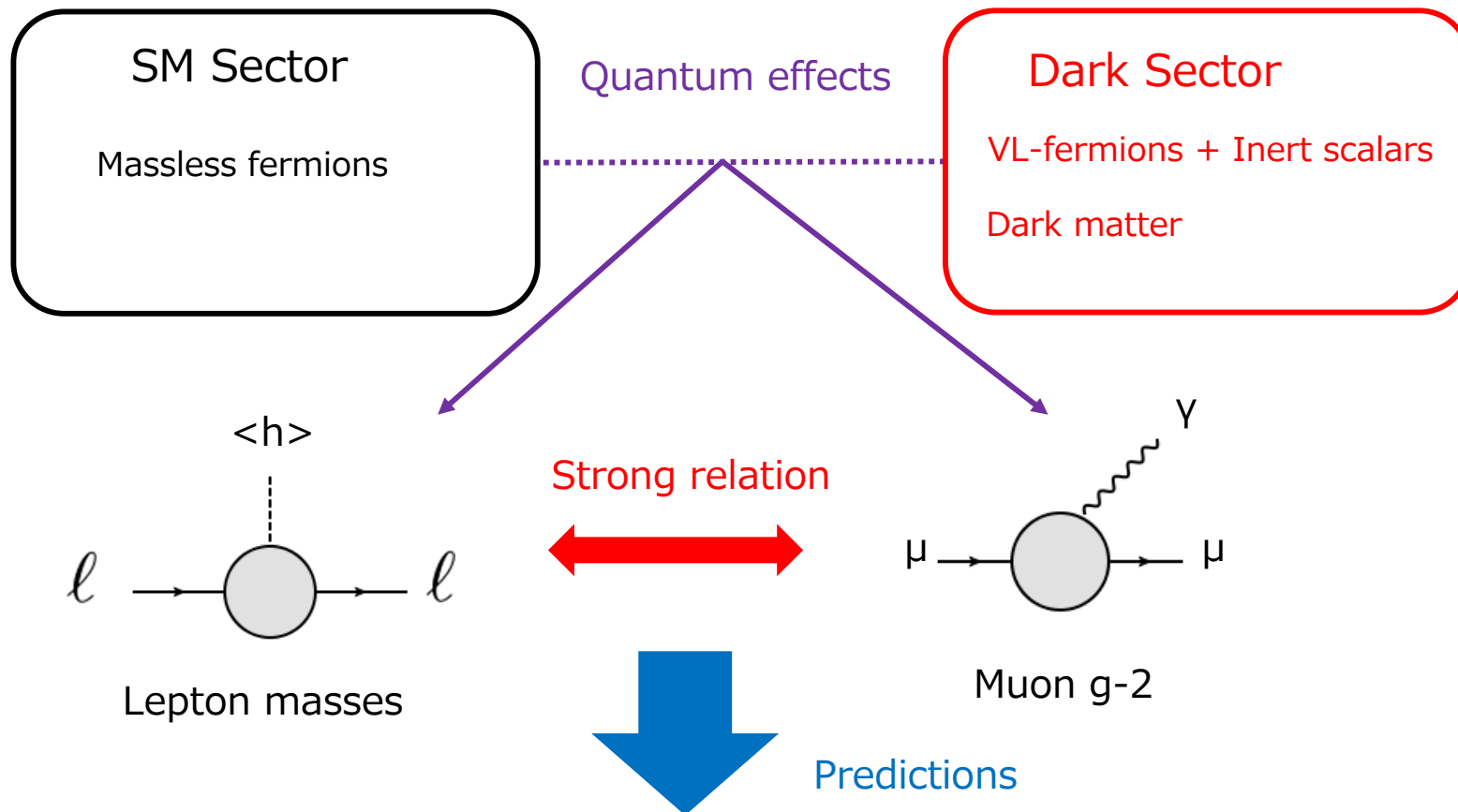
$$|\mathcal{M}_t|^2 \propto f_R^4 \quad (\text{for } \theta \sim \pi/2)$$



- New fermion can sizably change the cross section.
- Polarized beam would be useful to extract the signal.

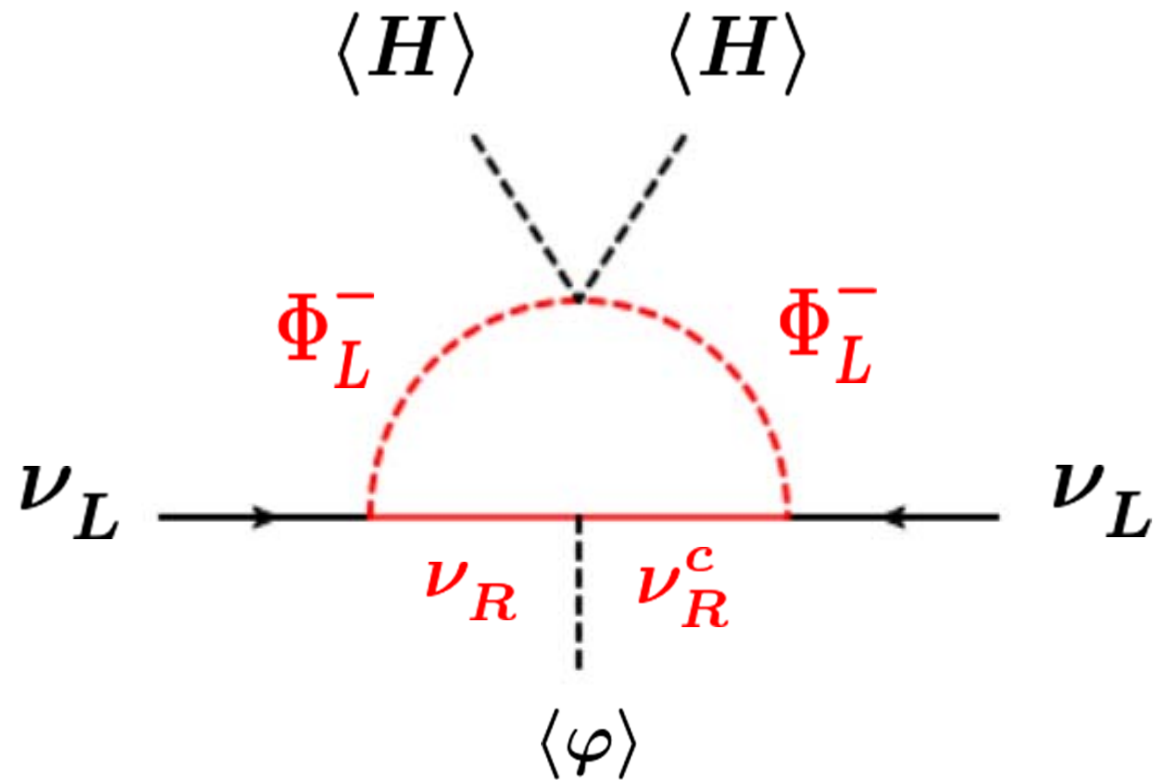
# Summary

Radiative charged seesaw scenarios can naturally solve DM and  $(g-2)_\mu$ .



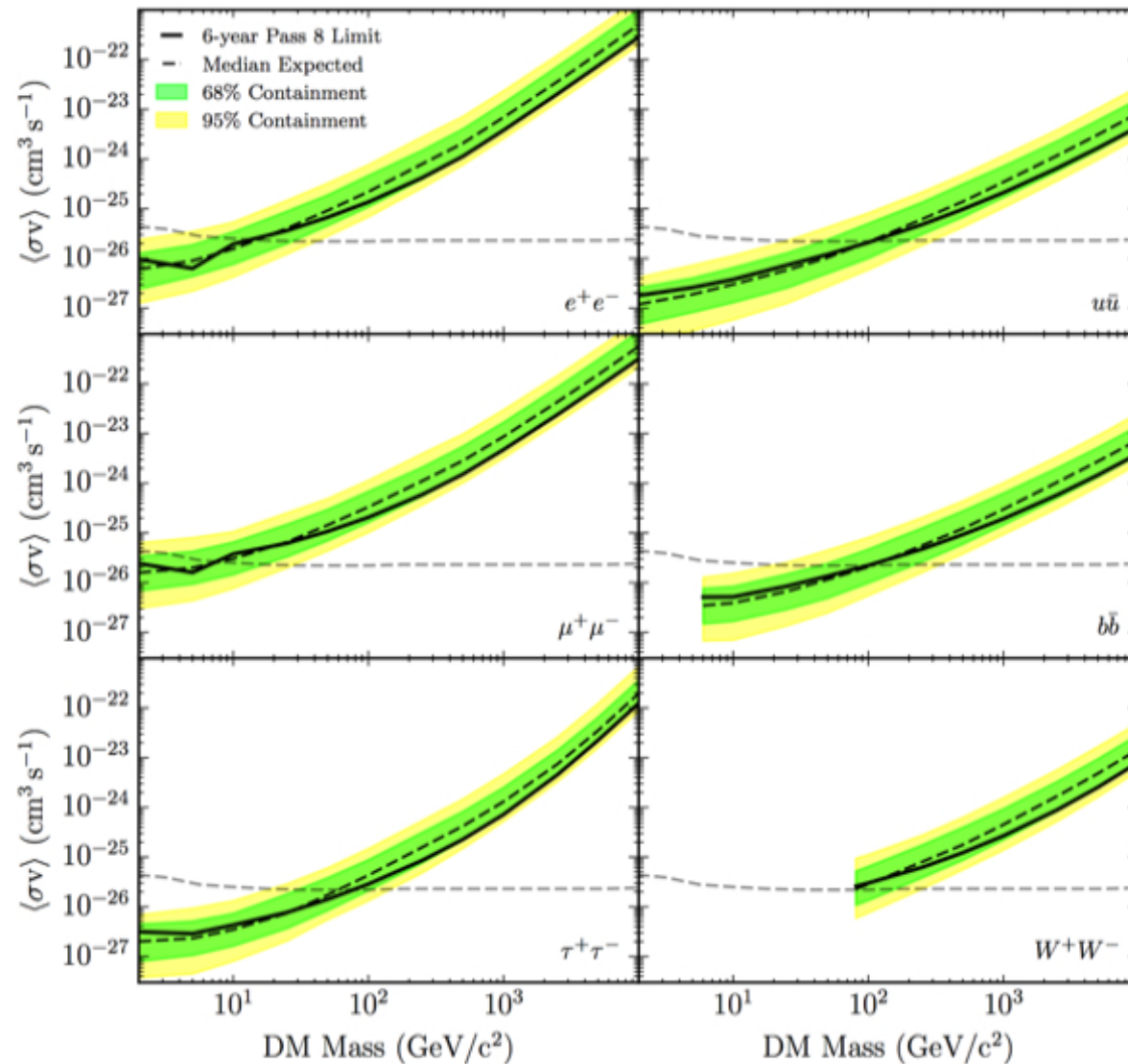
- (1) Large deviations in the **muon Yukawa coupling**
- (2) **Light dark scalars** can directly be detected at collider experiments.

# Neutrino masses



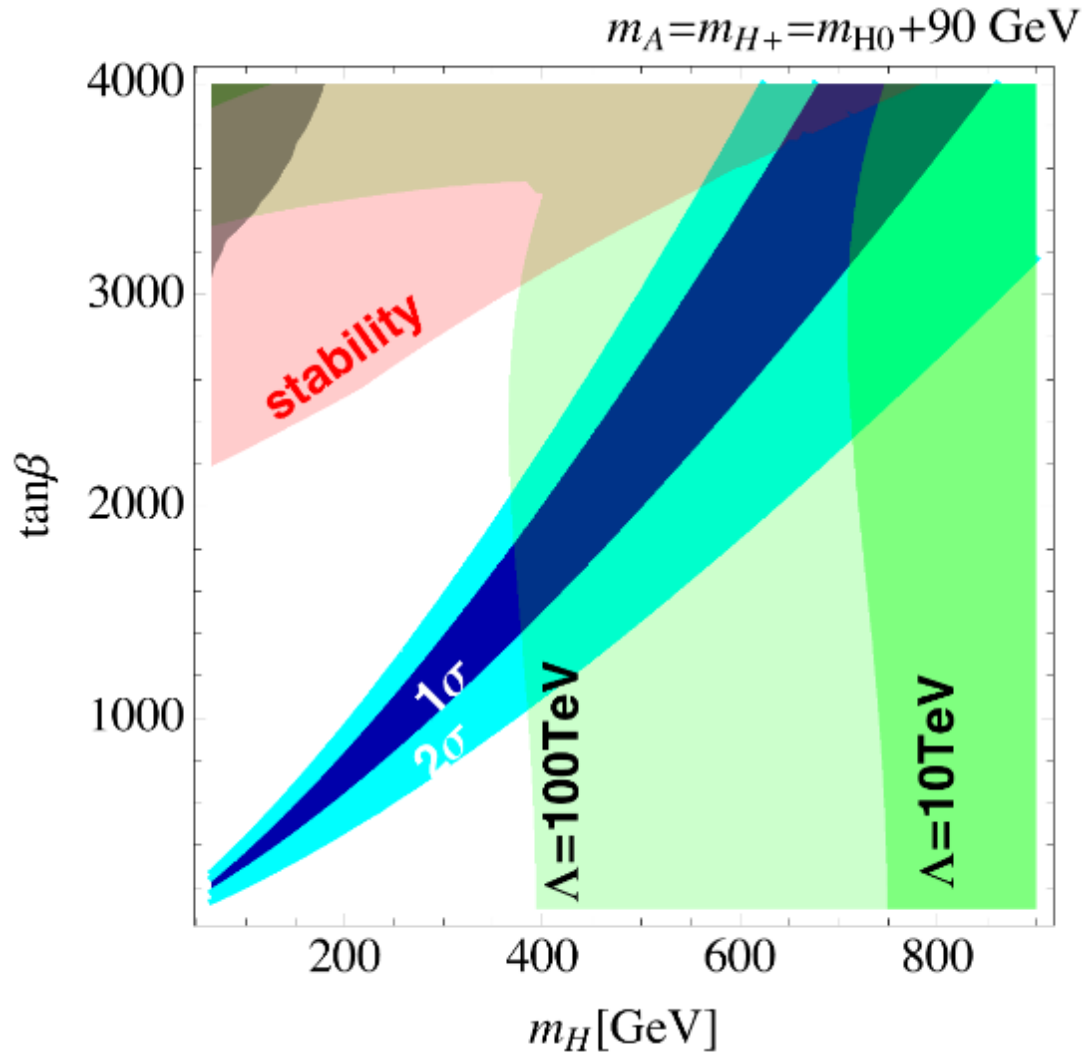
# Dark matter indirect searches

1503.02641 (Fermi-LAT)



The limits on the thermally averaged annihilation cross section of dark matter as a function of energy. The different graphs represent various annihilation channels. (Credit: [Fermi-LAT collab](#))





$m_{H^0} [\text{GeV}]$	$m_{A^0}(=m_{H^\pm}) [\text{GeV}]$	$\tan \beta$	$\sigma_{13\text{TeV}} [\text{fb}]$	$N_{\mu\text{-THDM}}$	$\mathcal{L}_{3\sigma} [\text{fb}^{-1}]$
600	700	3000	0.41	6.6	-
620	710	3000	0.369	5.9	-
640	730	3100	0.316	5.2	44
660	750	3300	0.2707	4.5	58
680	770	3400	0.2334	3.9	75
700	790	3700	0.20	3.4	97

# New physics contribution at 1-loop

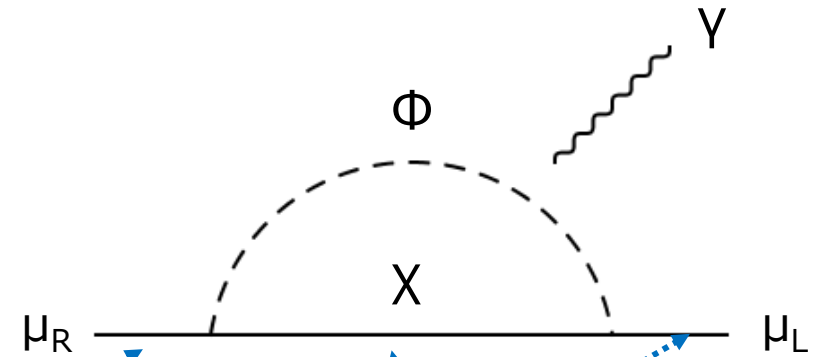
- Effective dim. 5 dipole operator

$$\mathcal{L}_{\text{eff}} = -\frac{1}{2} \left( \frac{e}{2m} a_f \right) \bar{\mu}_L \sigma^{\mu\nu} \mu_R F_{\mu\nu} + \text{h.c.}$$

- New interactions:

$$\mathcal{L}_{\text{new}} = \bar{\chi} (f_L P_L + f_R P_R) \mu \phi + \text{h.c.}$$

$$a_\mu = -\frac{Q_\chi}{16\pi^2} \frac{m_\mu}{M_\chi} \left[ f_L^2 \frac{m_\mu}{2M_\chi} G_2 \left( \frac{m_\phi^2}{M_\chi^2} \right) + f_R^2 \frac{m_\mu}{2M_\chi} G_2 \left( \frac{m_\phi^2}{M_\chi^2} \right) + f_L f_R G_1 \left( \frac{m_\phi^2}{M_\chi^2} \right) \right]$$



For  $M_\chi \gg m_\mu$ , the last term can be dominant. In this case, we can estimate

$$a_\mu^{\text{NP}} \simeq 3 \times 10^{-9} \times \left( \frac{2 \text{ TeV}}{M_\chi} \right) \times \left( \frac{f_L}{0.1} \right) \times \left( \frac{f_R}{0.1} \right)$$